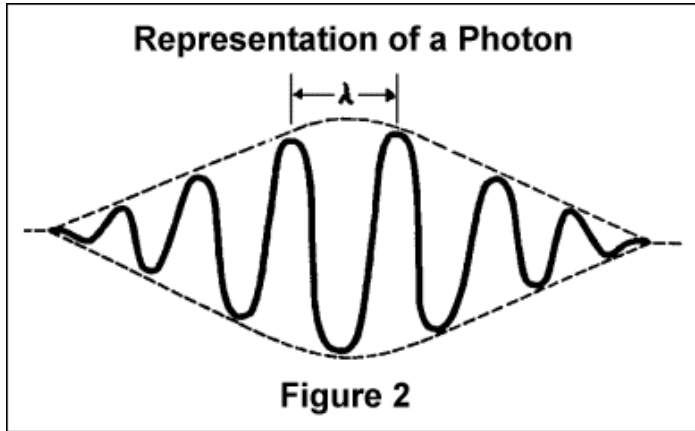
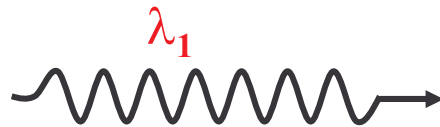


# אפקט קומפטון

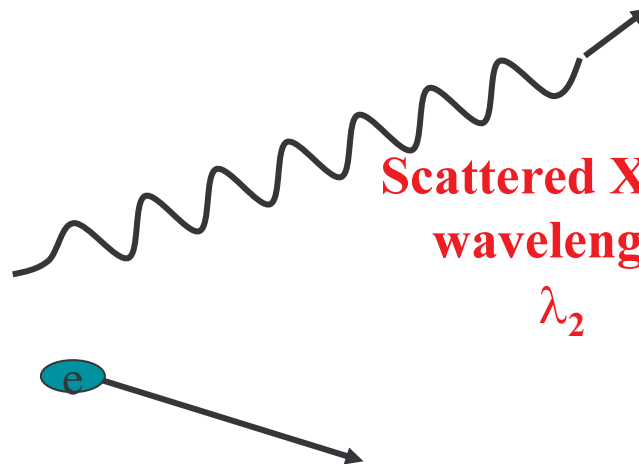


Incident X-ray  
wavelength



M  
A  
T  
T  
E  
R

Scattered X-ray  
wavelength



$$\lambda_2 > \lambda_1$$

$$\lambda_{\text{compton}} = \frac{h}{mc}$$

Electron comes flying out

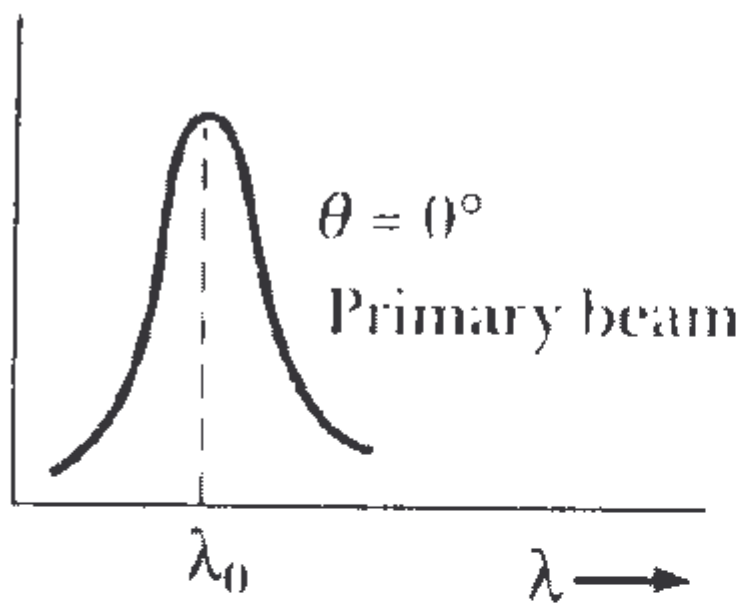
$$\lambda_{\text{compton}}(e) = 0.024 \text{ \AA}$$



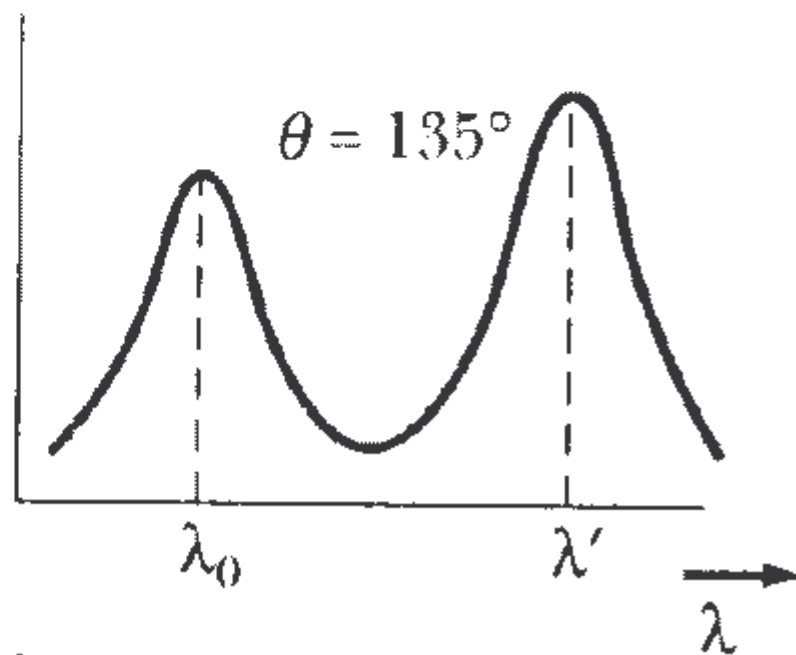
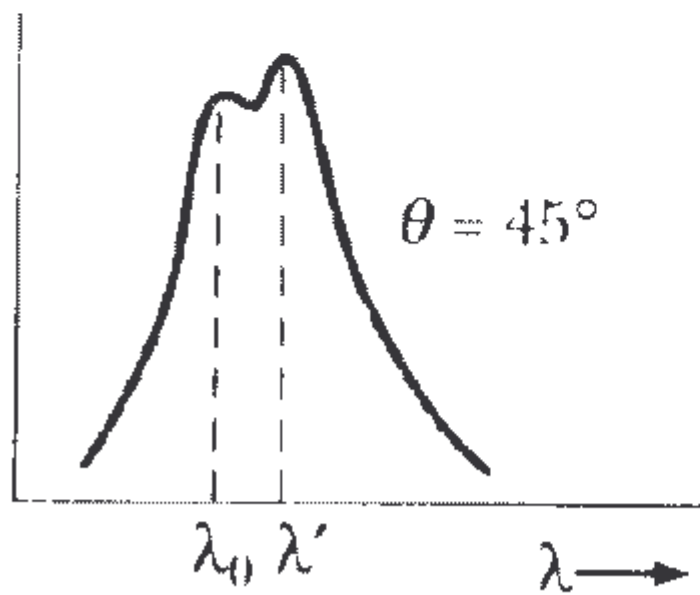
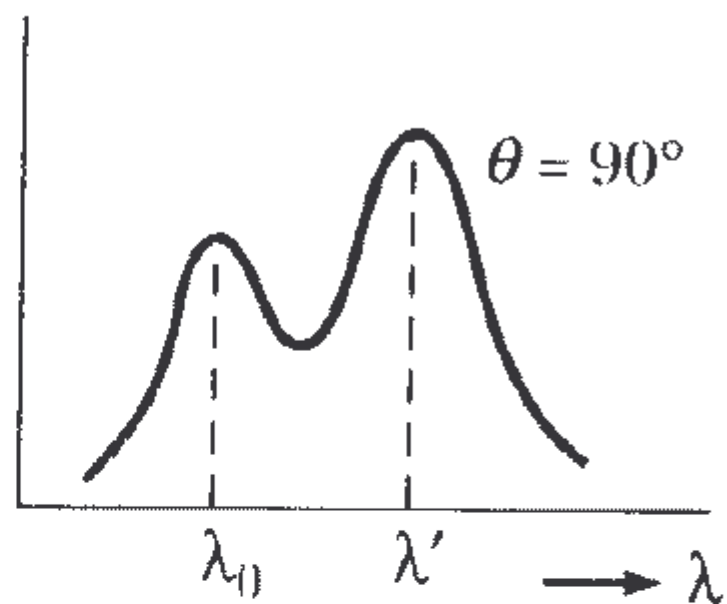
קומפטון (1892-1962)

Nobel prize, 1927,  
for his discovery of the  
effect named after him  
(performed in 1922)

Intensity



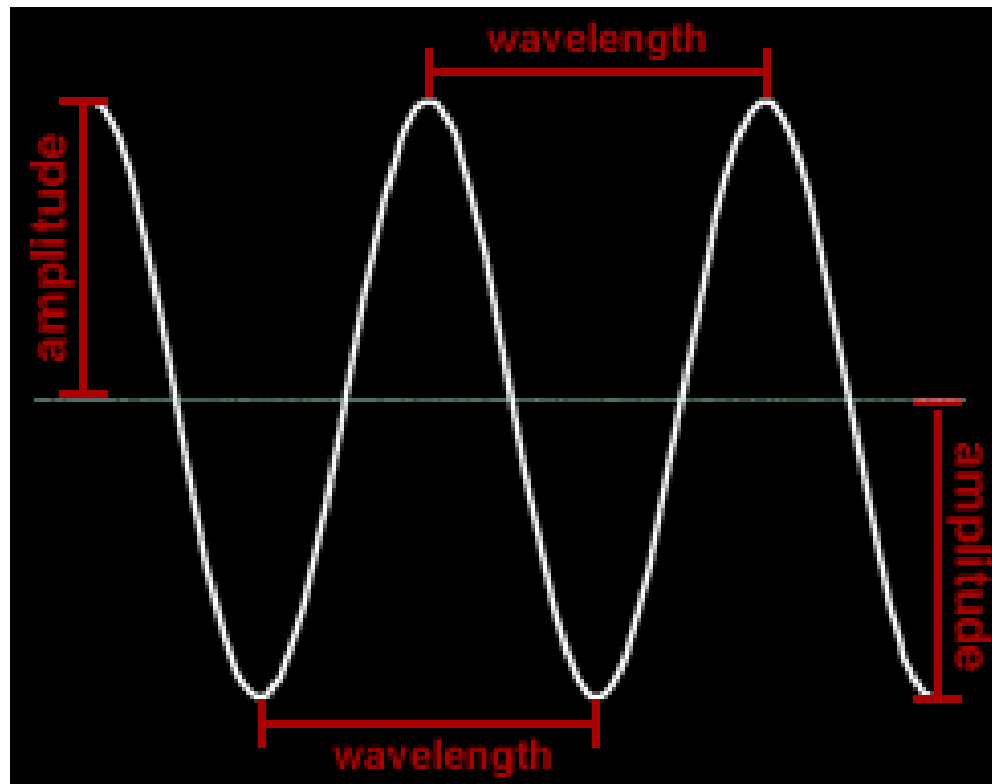
Intensity



(b)

# Light Waves

Until about 1900, the classical wave theory of light described most observed phenomenon.



Light waves:

Characterized by:

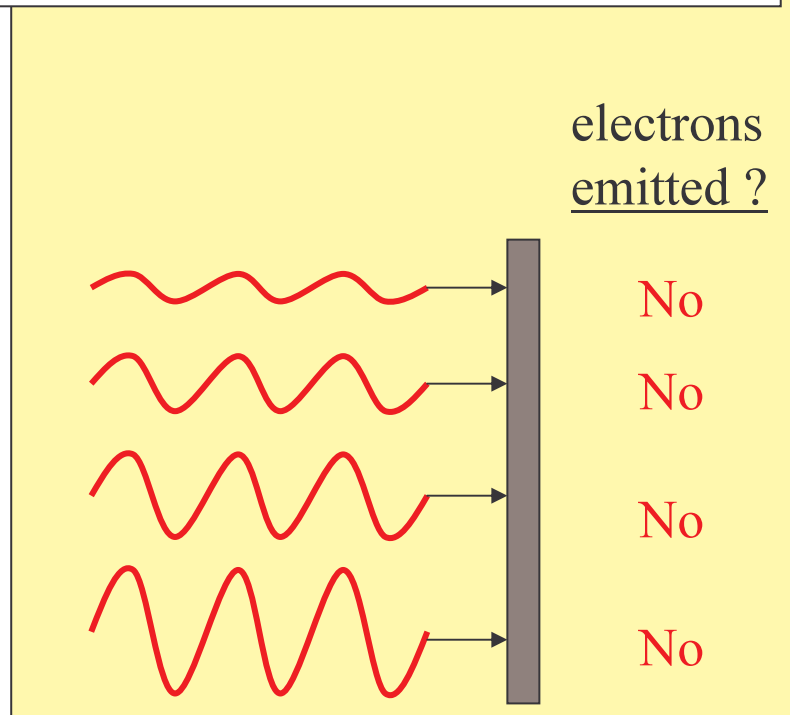
- Amplitude (A)
- Frequency ( $\nu$ )
- Wavelength ( $\lambda$ )

$$\text{Energy} \propto A^2$$

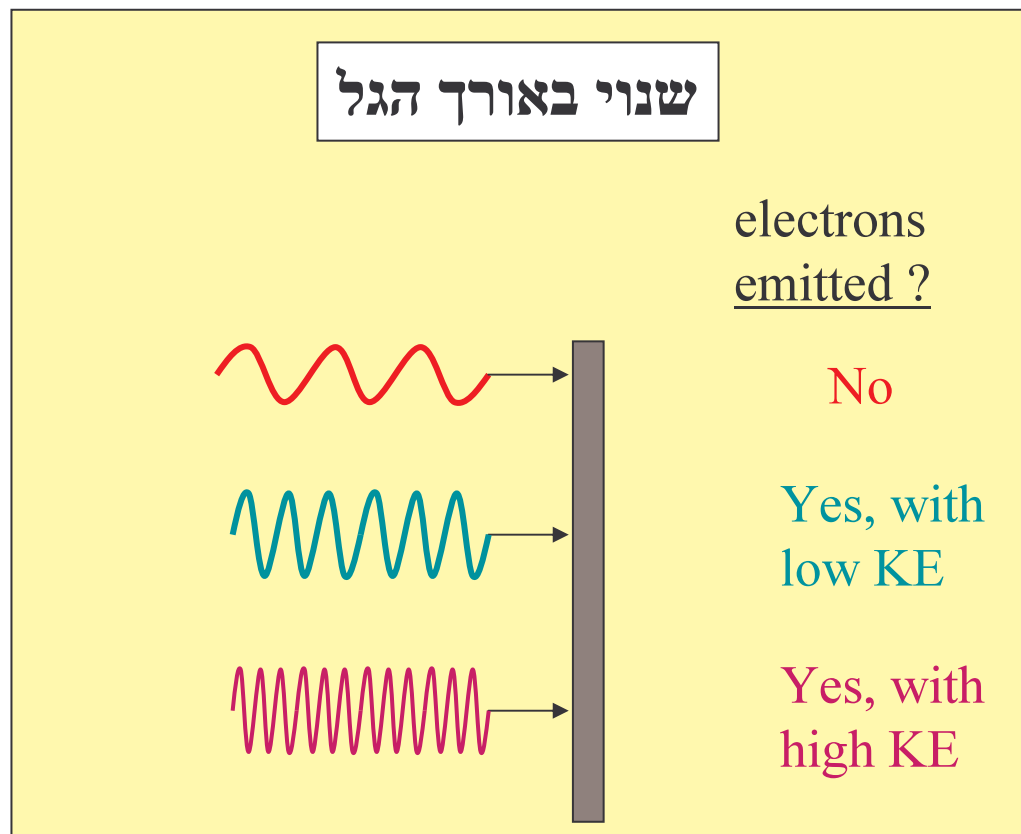
# האפקט הפוטואלקטרי

## השיטה הקלאסית

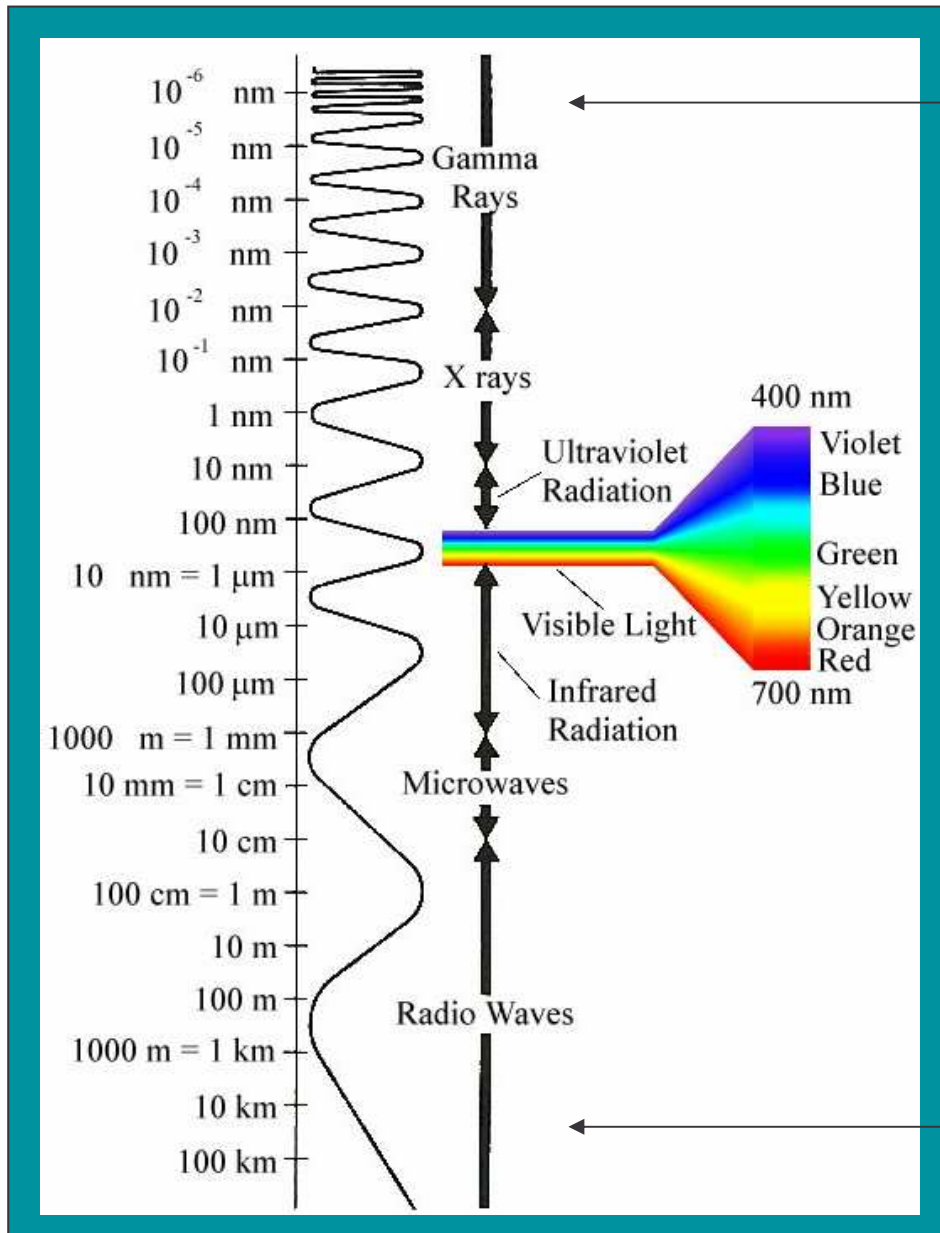
הגדלת האנרגיה על ידי עצמת הגל



שנוי באורך הגל



# The Electromagnetic Spectrum



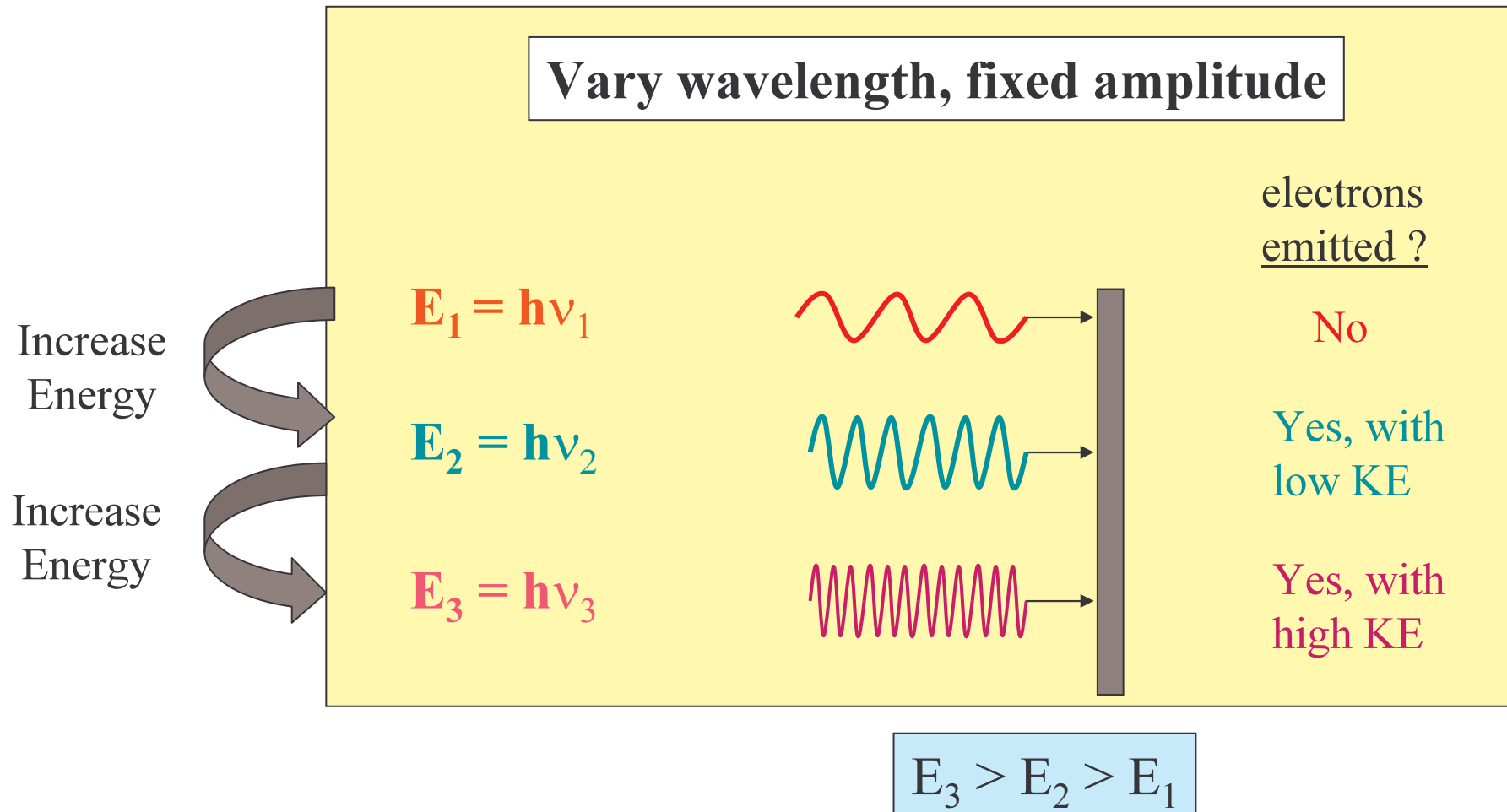
Shortest wavelengths  
(Most energetic photons)

$$E = h\nu = hc/\lambda$$

$h = 6.6 \times 10^{-27}$  [erg\*sec]  
(Planck's constant)

Longest wavelengths  
(Least energetic photons)

# Interpretation of Photoelectric Effect



**Photoelectric Effect Applet**

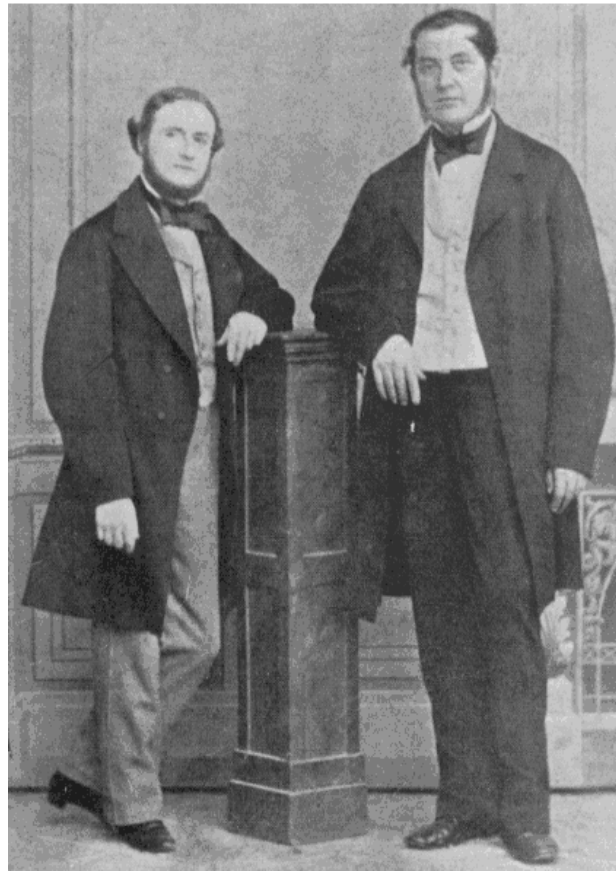


רוברט מיליקן (1868-1953)

Nobel Prize, 1923,

for his work on the elementary charge of electricity and on the photoelectric effect.



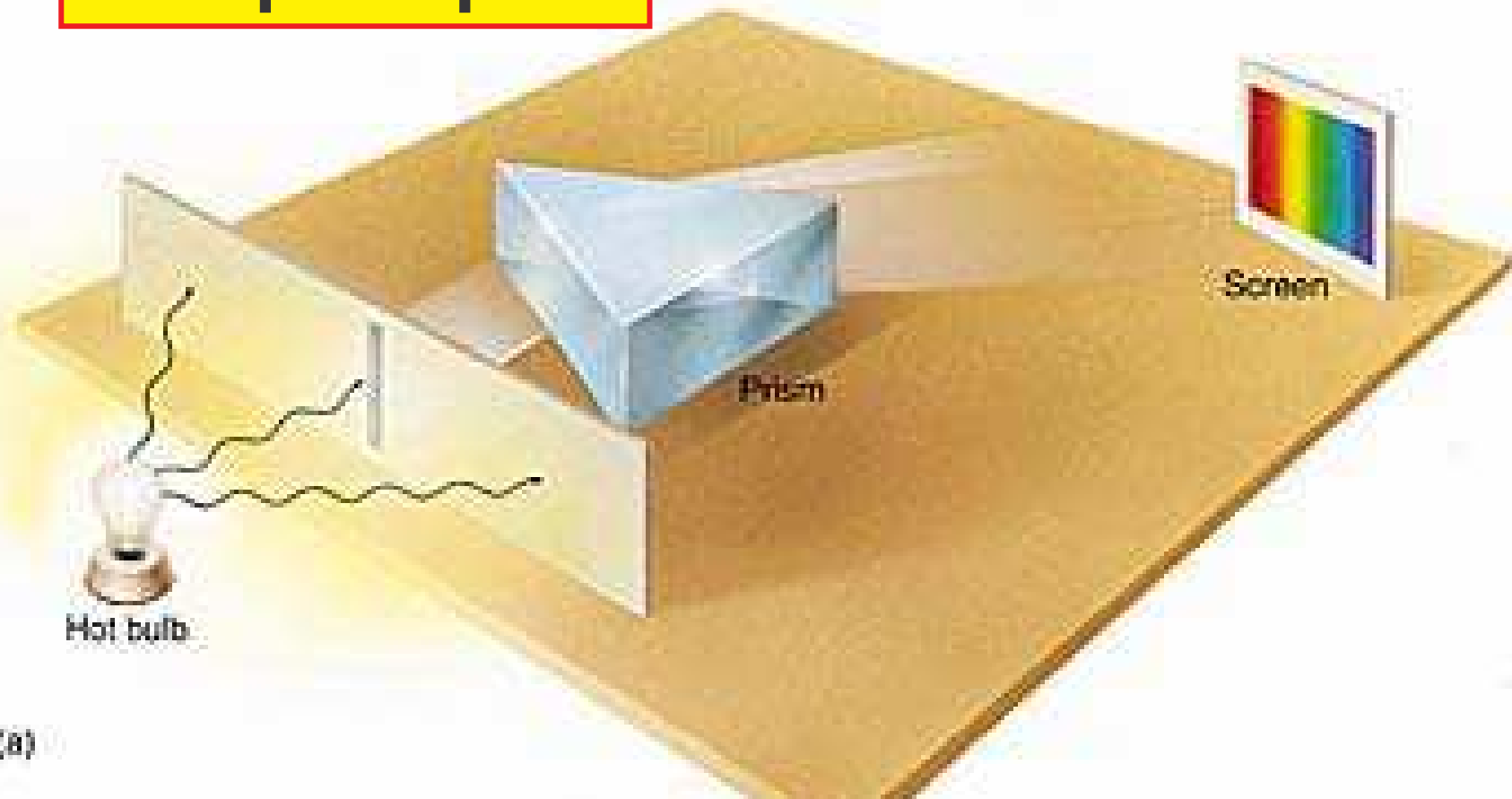


קירכהוף (1824-1887)

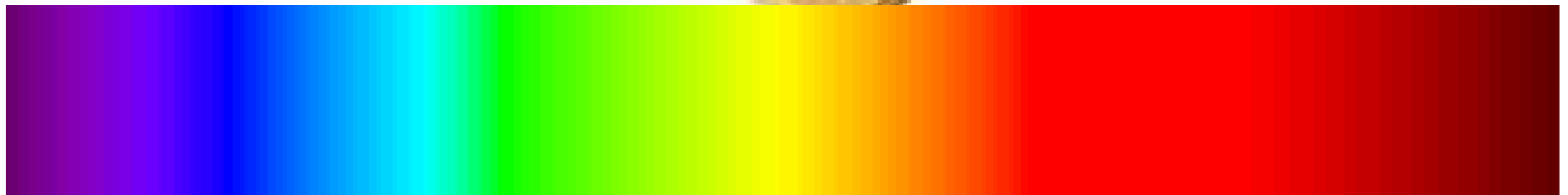


רוברט בונזן (1811-1899)

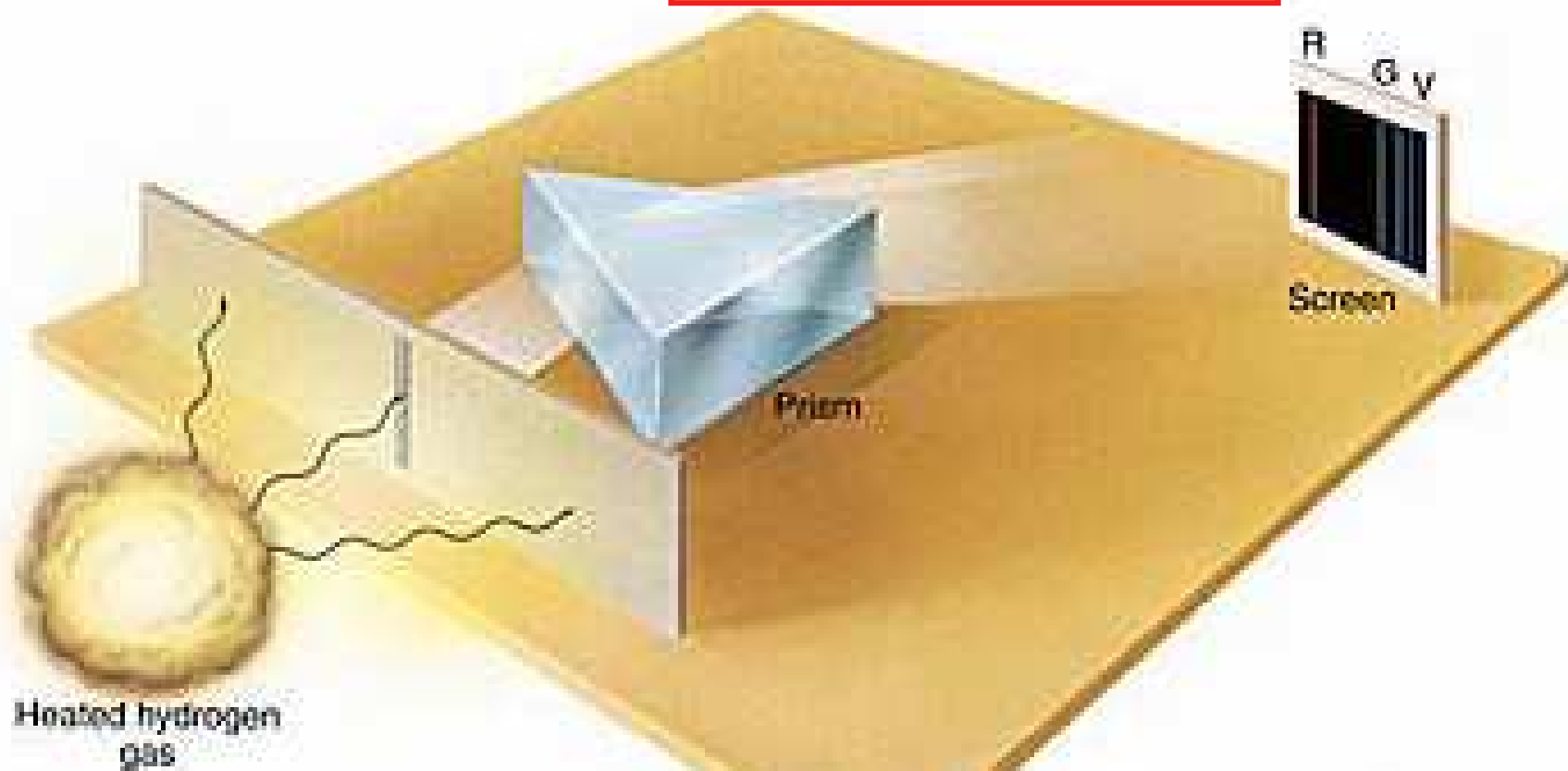
# ספקטרוסקופיה



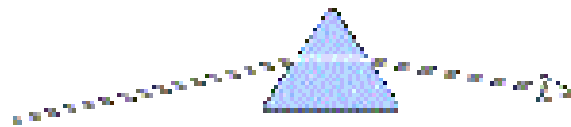
(a)



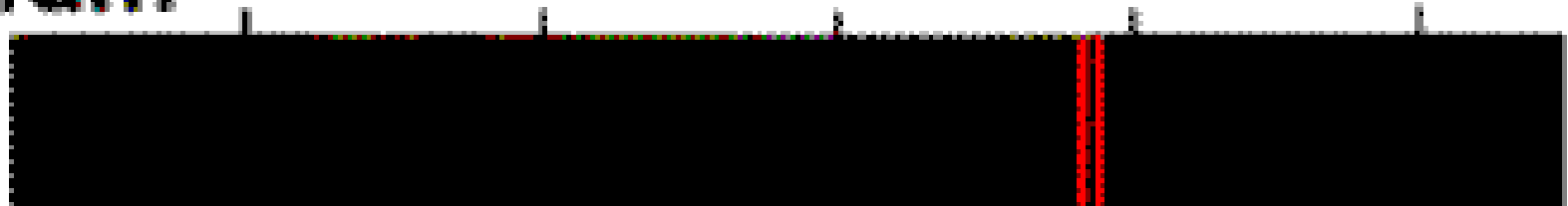
# ספקטרוסקופיה



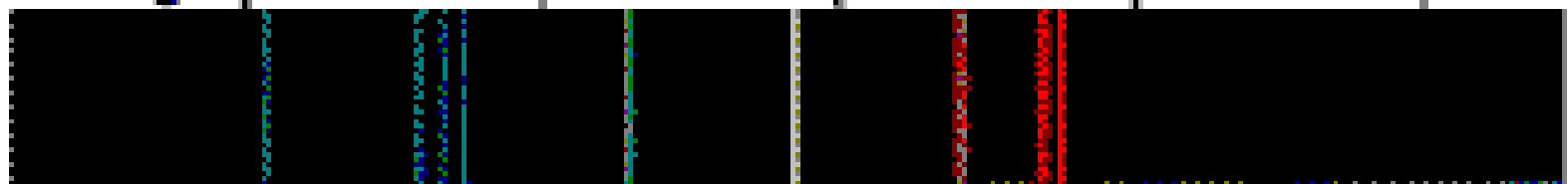
Hot Gas



Sodium



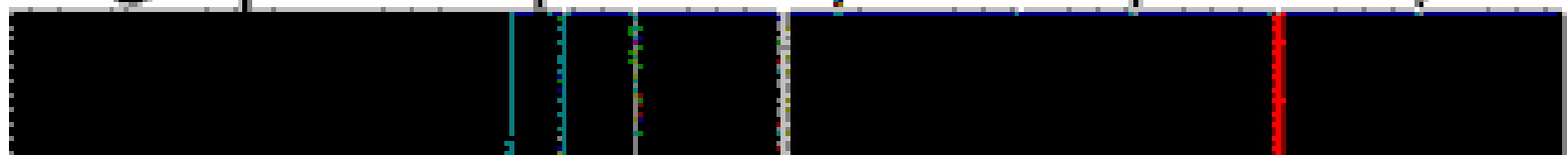
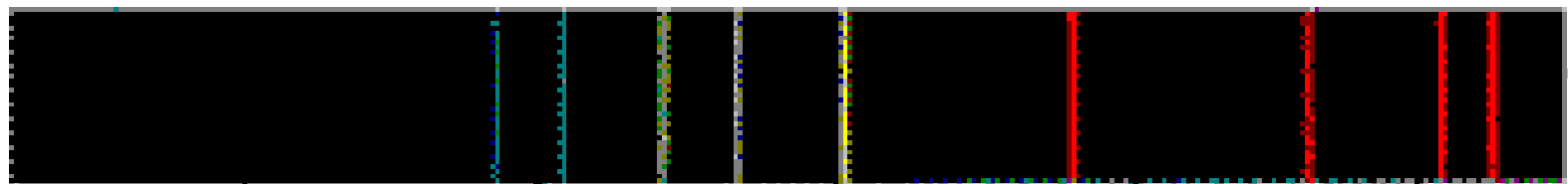
Mercury

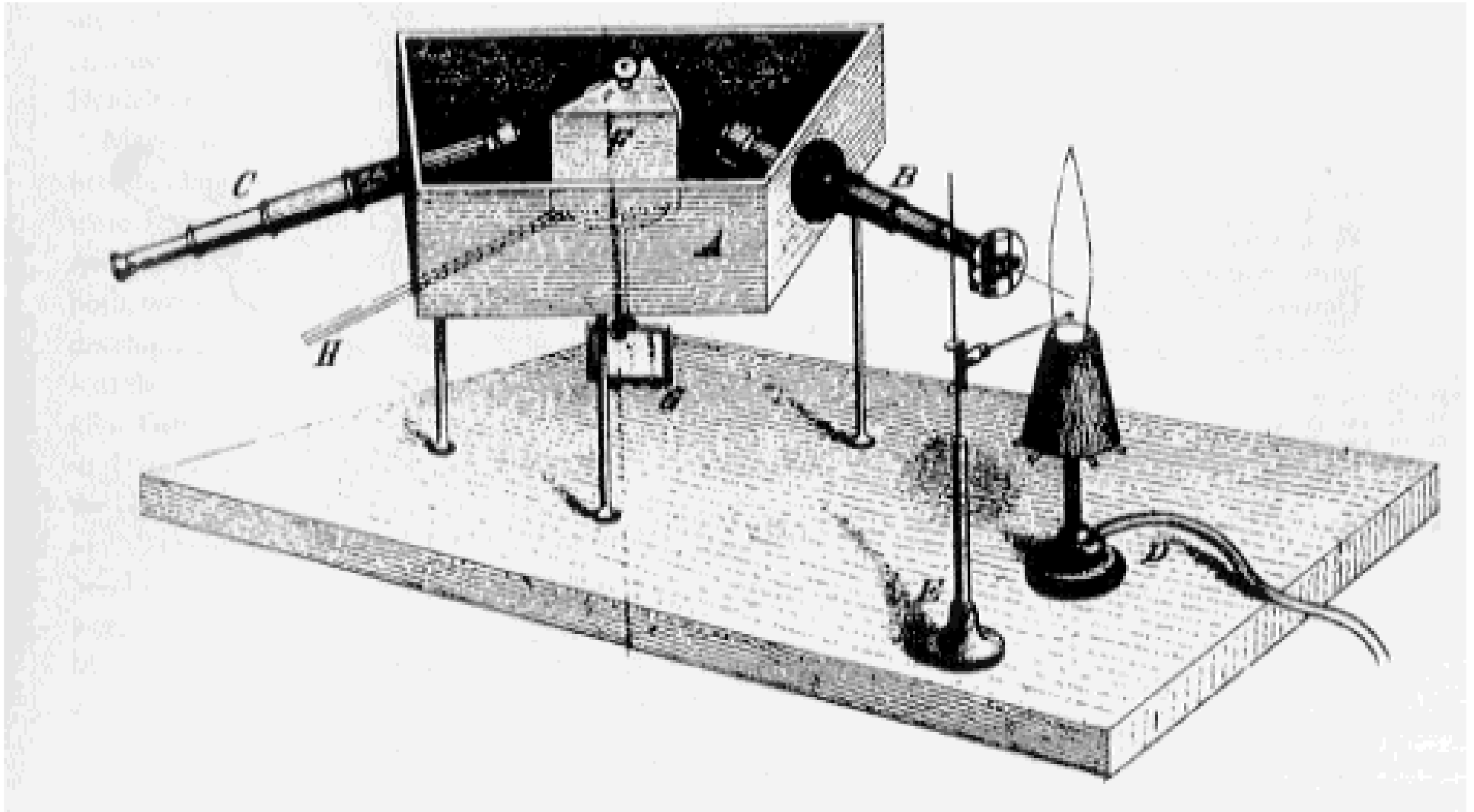


Helium

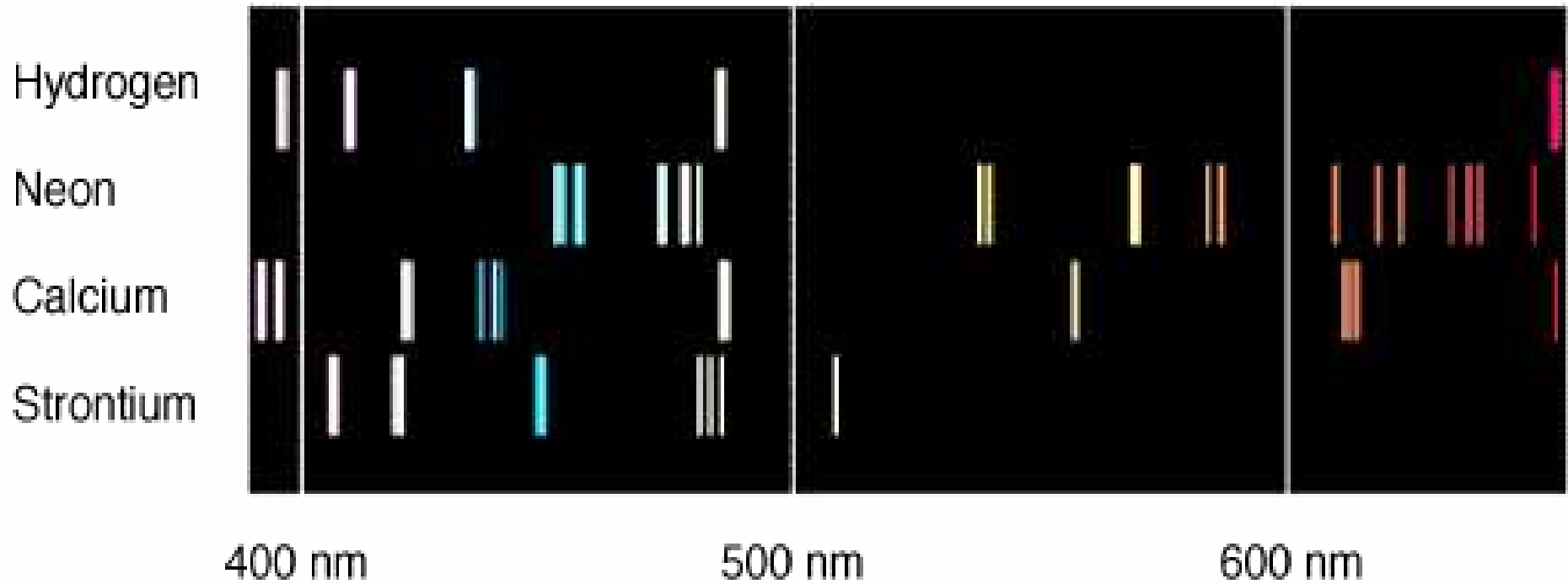


Hydrogen





# ספקטרה של פליטה



# Periodic Table of the Elements

	IA															0		
1	1 H	IIA															2 He	
2	3 Li	4 Be										5 B	6 C	7 N	8 O	9 F	10 Ne	
3	11 Na	12 Mg	IIIB	IVB	VB	VIB	VII B	VIII			IB	IB	13 Al	14 Si	15 P	16 S	17 Cl	18 Ar
4	19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr
5	37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe
6	55 Cs	56 Ba	*La	72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn
7	87 Fr	88 Ra	+Ac	104 Rf	105 Ha	106 Sg	107 Ns	108 Hs	109 Mt	110	111	112	113					

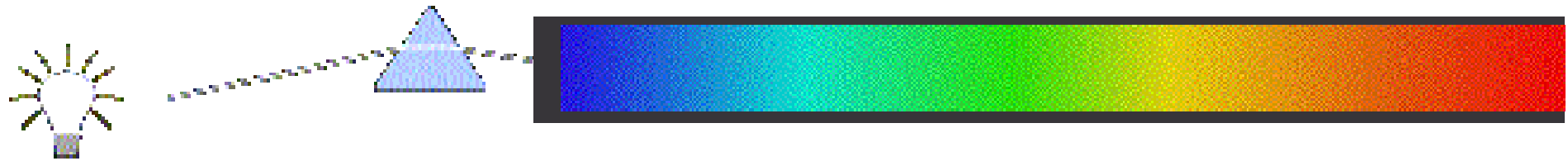
\* Lanthanide Series

58	59	60	61	62	63	64	65	66	67	68	69	70	71
Ce	Pr	Nd	Pm	Sm	Eu	Gd	Tb	Dy	Ho	Er	Tm	Yb	Lu

+ Actinide Series

90	91	92	93	94	95	96	97	98	99	100	101	102	103
Th	Pa	U	Np	Pu	Am	Cm	Bk	Cf	Es	Fm	Md	No	Lr

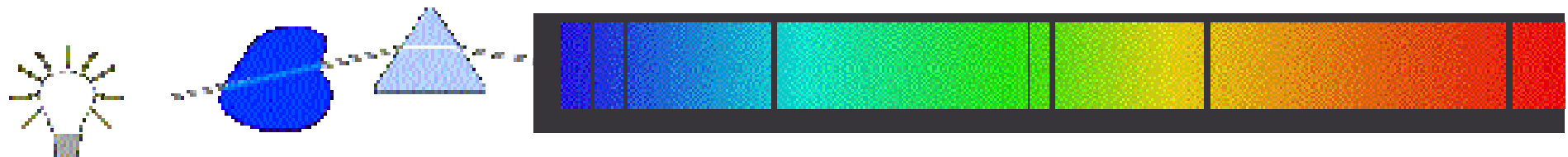
# ספקטרוסקופיה



Hot Gas



Cold Gas

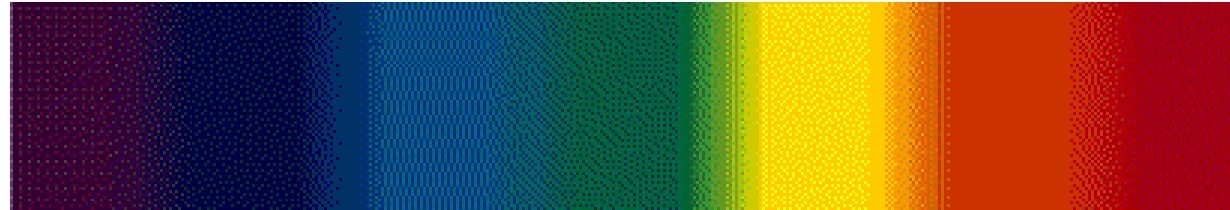




# שלושת סוגי הספקטרה

Continuous

A



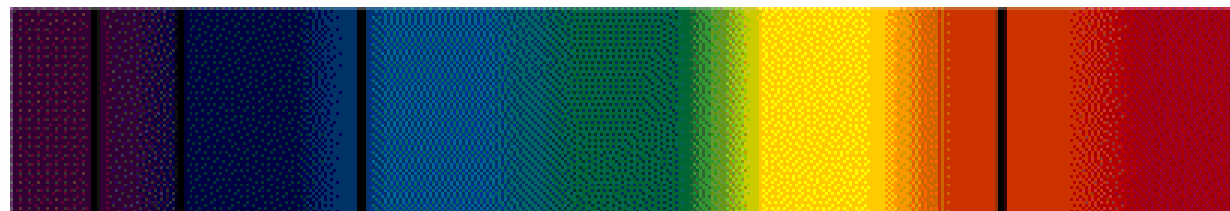
Emission line (hydrogen gas)

B

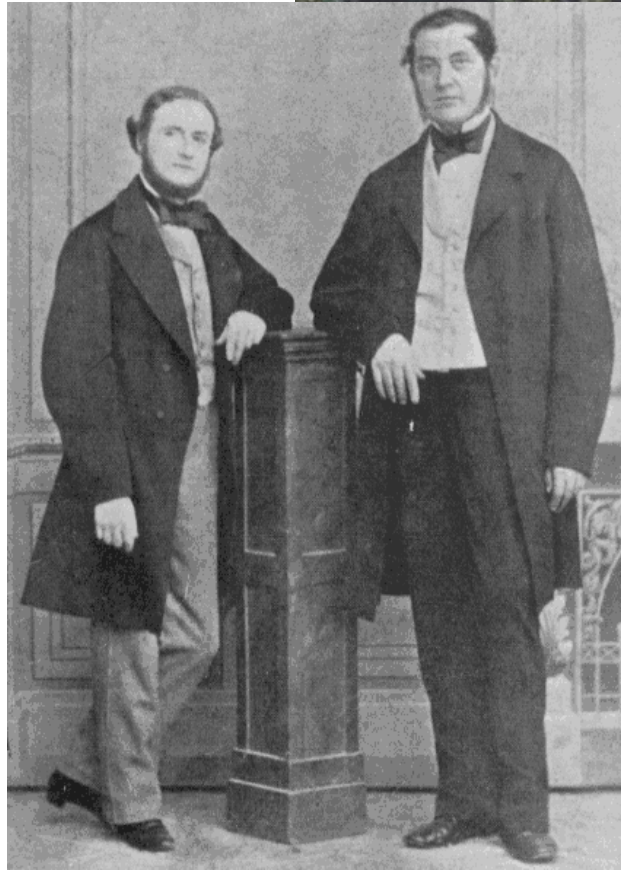


Absorption line (hydrogen gas)

C



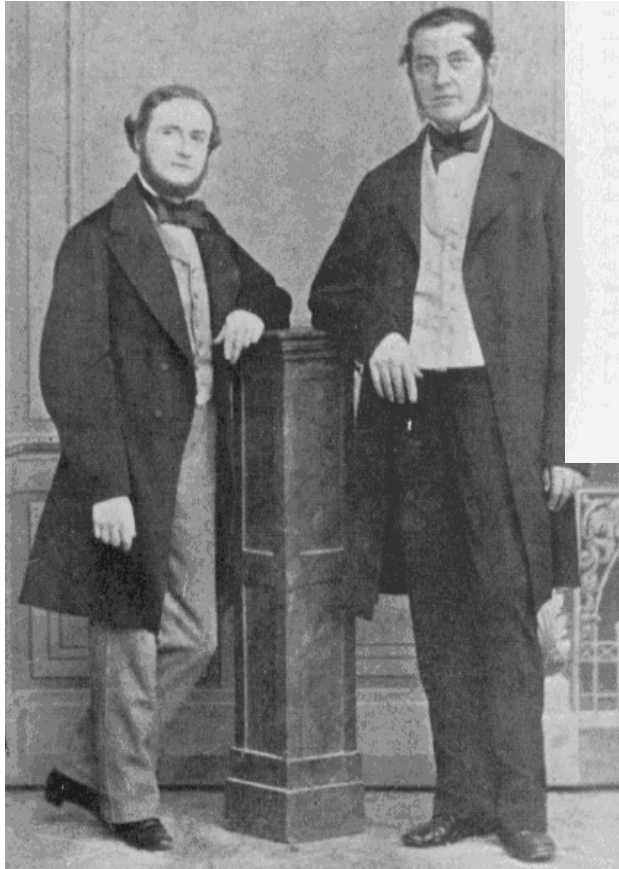
# היידלברג



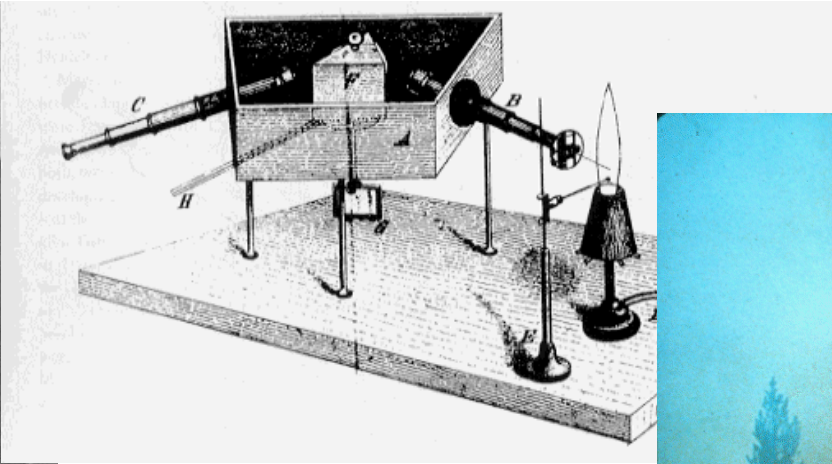
בונזן וקירכהוף



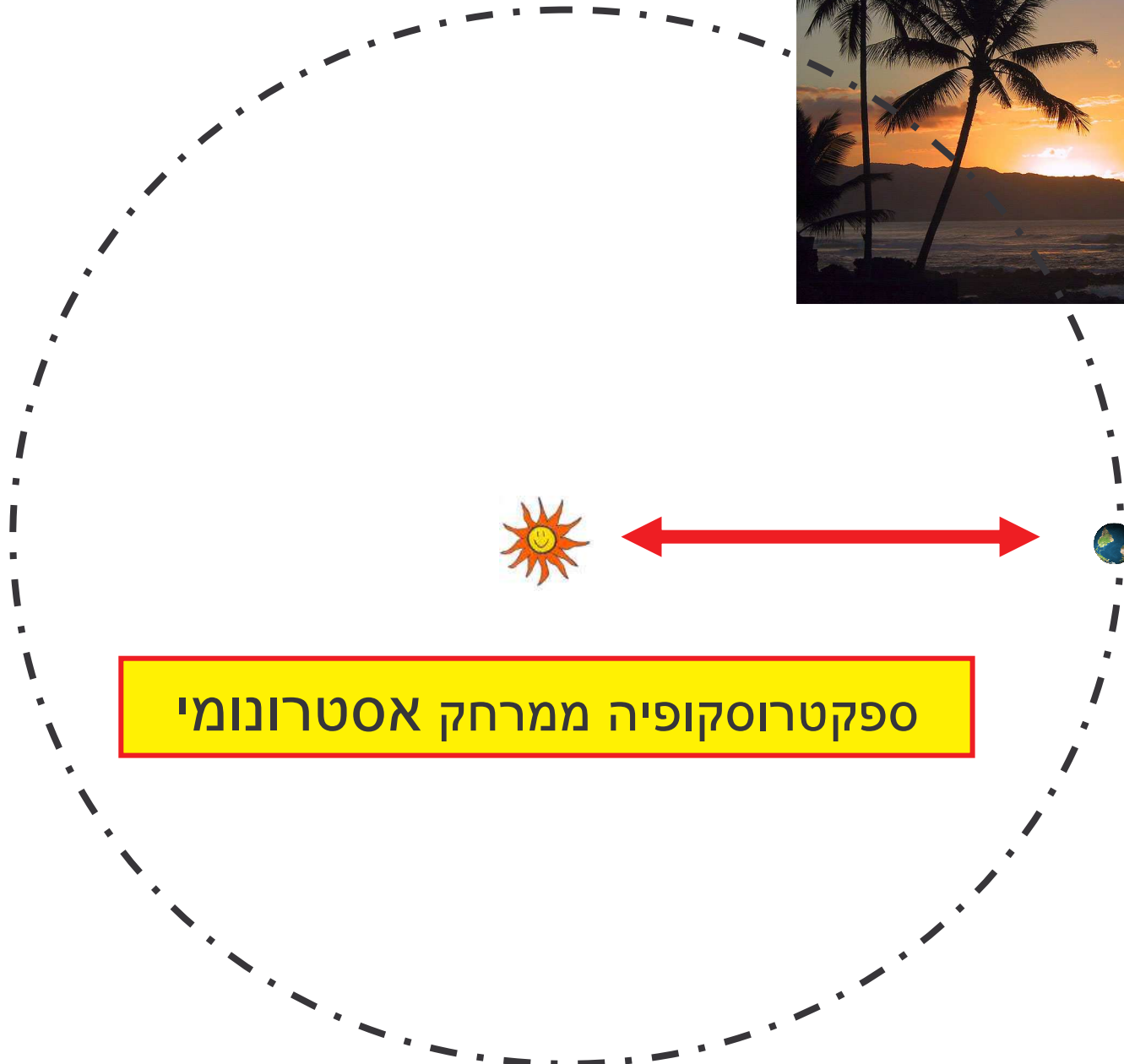




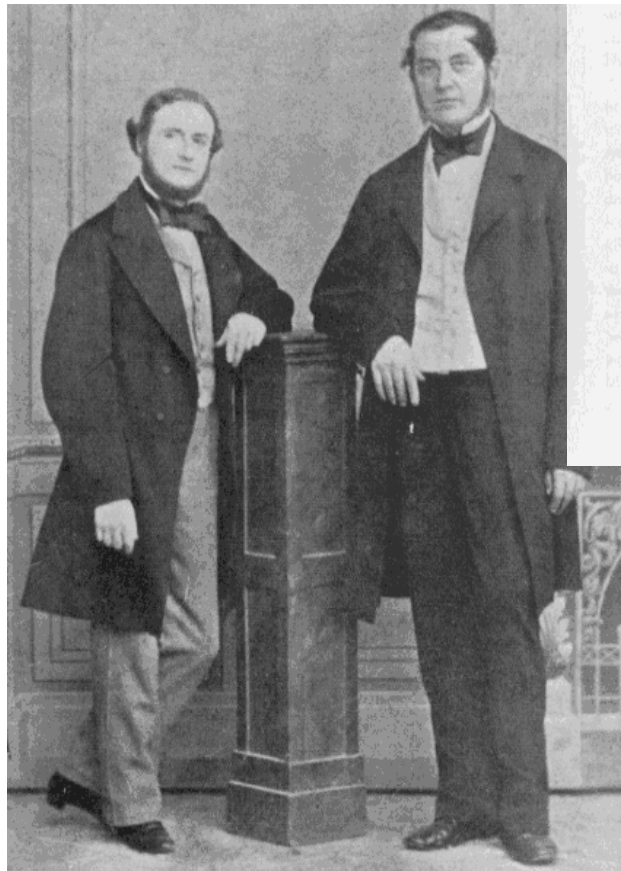
בונזן וקירכהוף



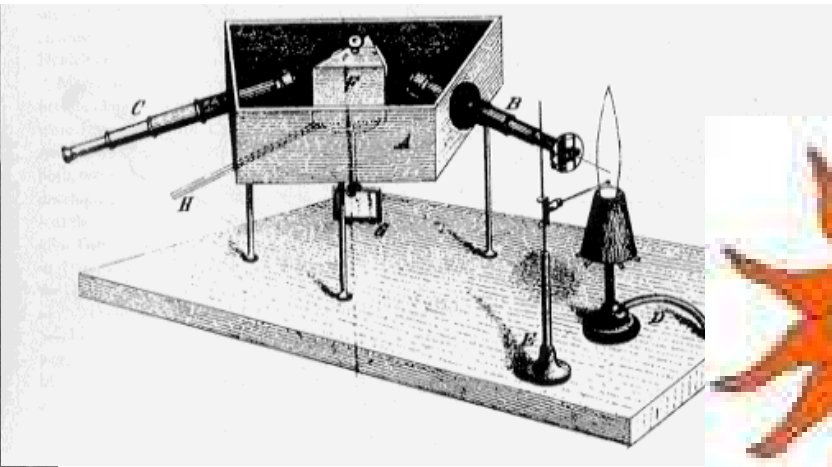
ספקטרוסקופיה ממרחק



ספקטרוסקופיה ממרחק אסטרונומי

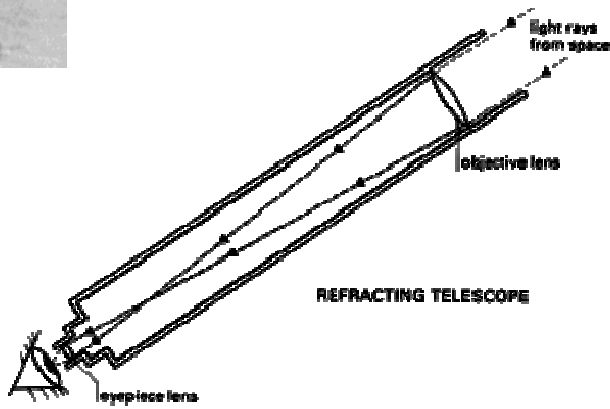


בונזן וקירכהוף

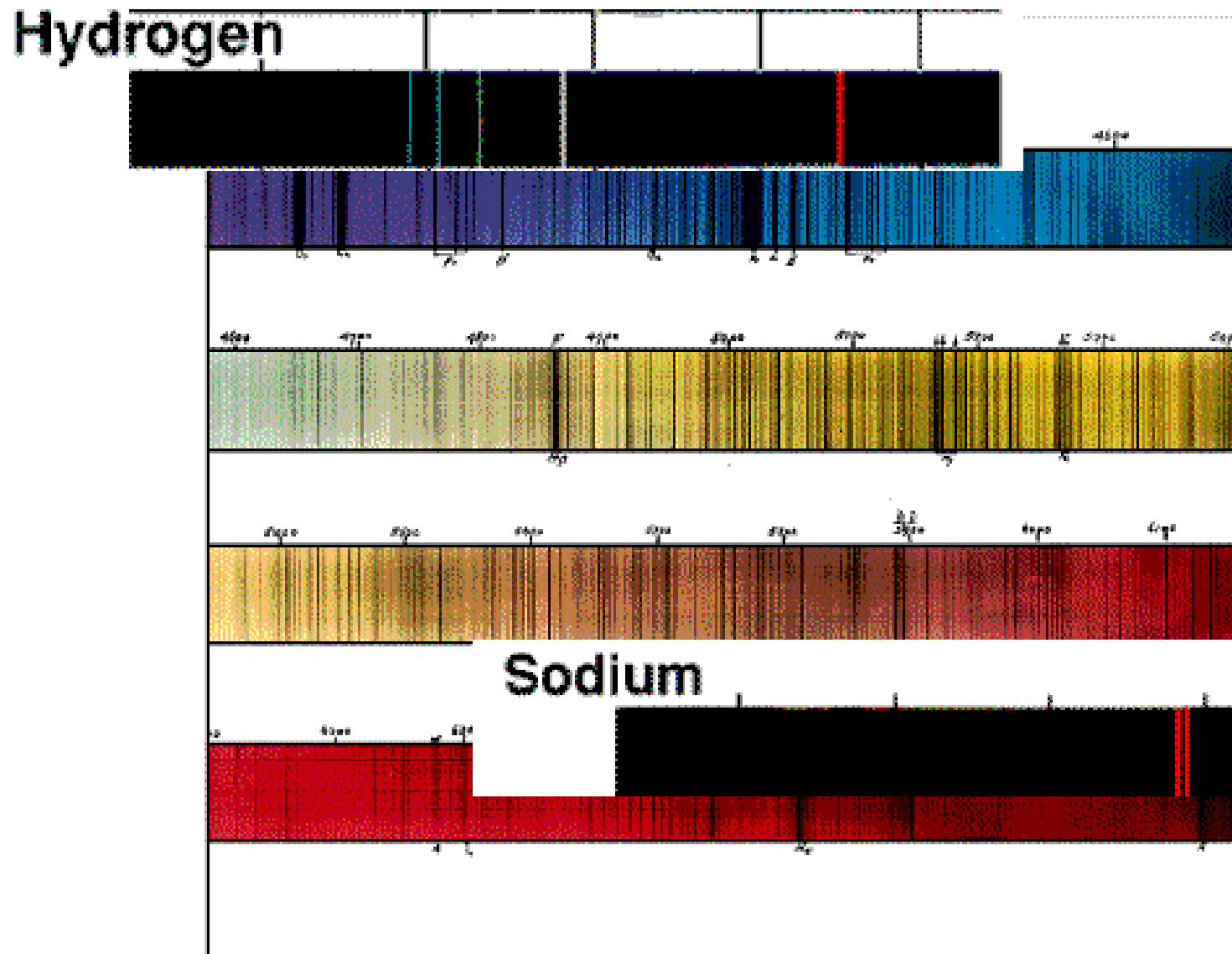


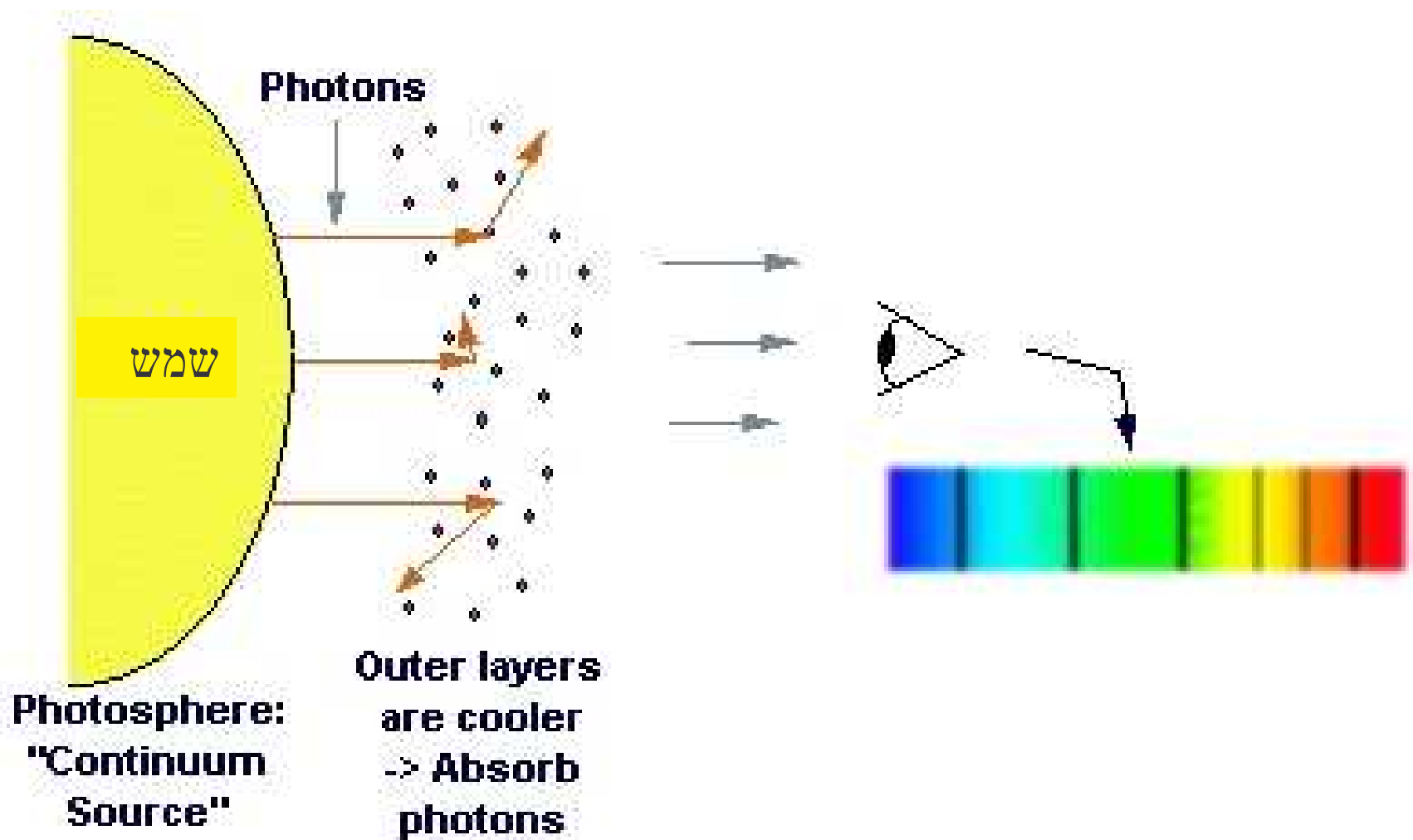
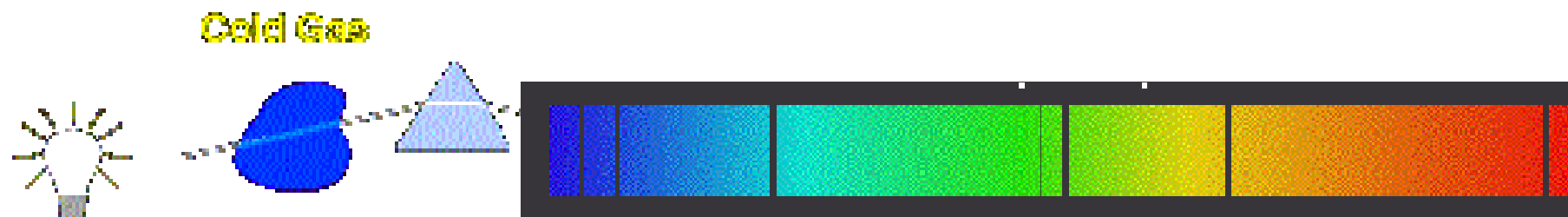
ספקטרוסקופיה אסטרונומית

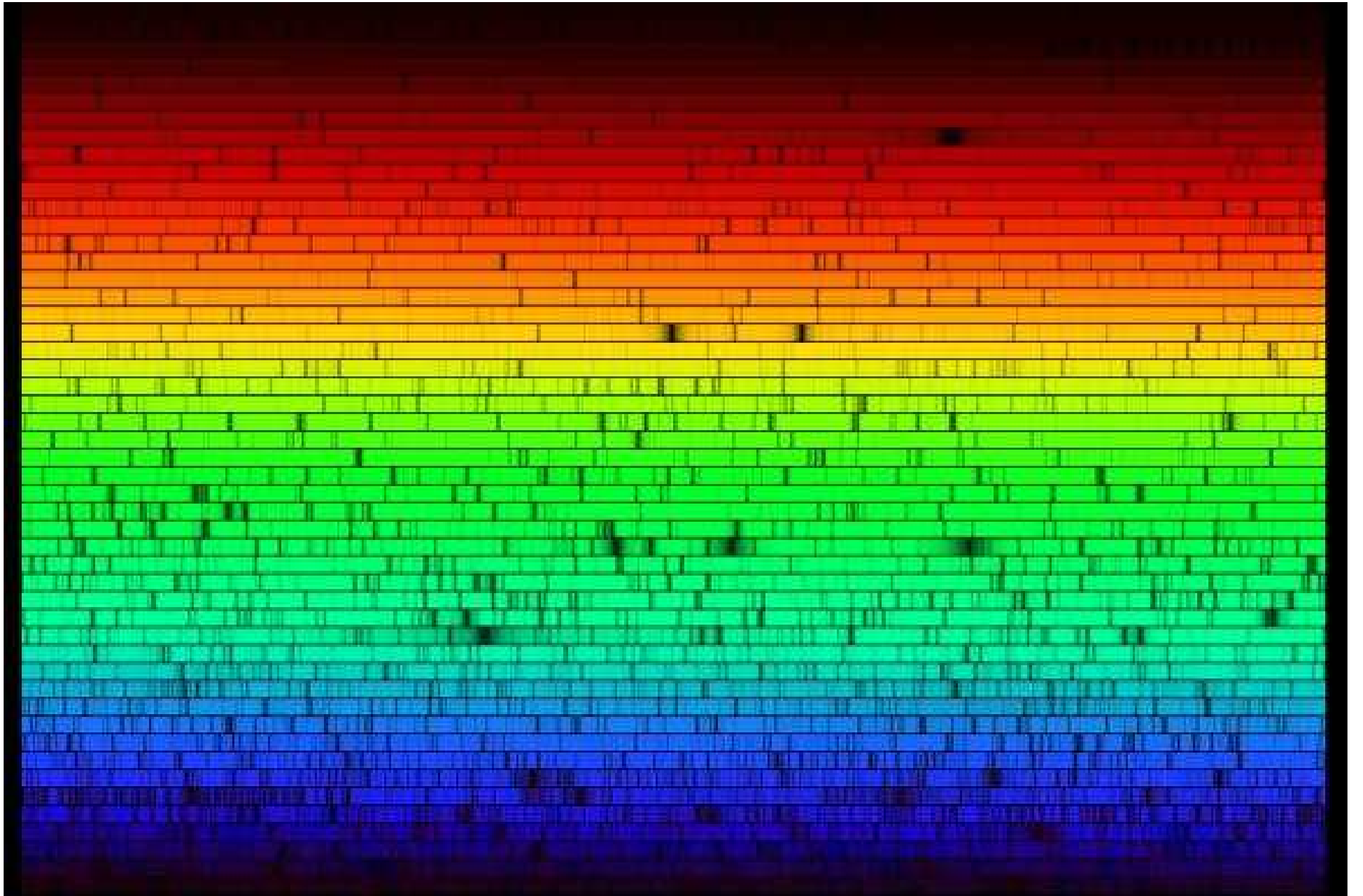
נפתחו השמיים  
וארא מראות אלהים



# קווי פראונאופר של השמש









# השמש עשויה:

מימן 91.2%

הליום 8.7%

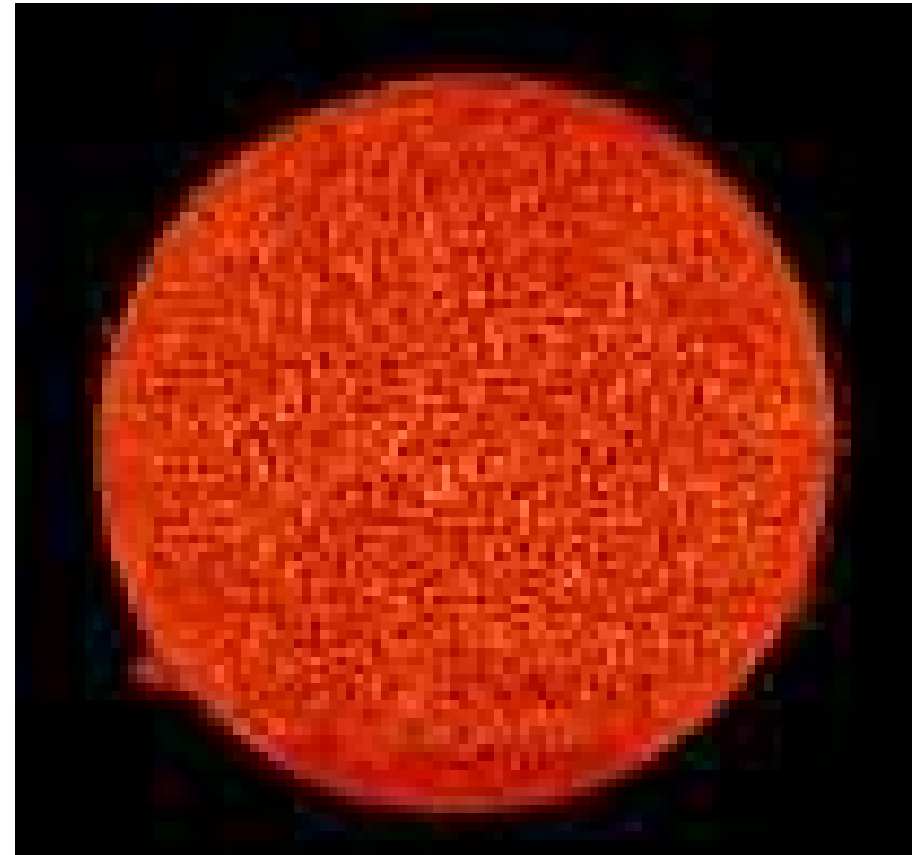
שאר היסודות 0.1%

טמפרטורת השפה – 6000 מעלות

## Periodic Table of the Elements

1A																	0									
1	2											3	4	5	6	7	8	9	10							
1	2	3	4											5	6	7	8	9	10							
H	He	Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar									
11	12	13	14	15	16	17	18	III B										IV B	V B	VI B	VII B	VIII	IX	X	XI	XII
19	20	21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36									
K	Ca	Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cu	Zn	Ga	Ge	As	Se	Br	Kr									
37	38	39	40	41	42	43	44	45	46	47	48	49	50	51	52	53	54									
Rb	Sr	Y	Zr	Nb	Mo	Tc	Ru	Rh	Pd	Ag	Cd	In	Sn	Sb	Te	I	Xe									
55	56	57	72	73	74	75	76	77	78	79	80	81	82	83	84	85	86									
Cs	Ba	*La	Hf	Ta	W	Re	Os	Ir	Pt	Au	Hg	Tl	Pb	Bi	Po	At	Rn									
87	88	89	104	105	106	107	108	109	110	111	112	113														
Fr	Ra	+Ac	Rf	Ha	Sg	Ns	Hs	Mt	110	111	112	113														

## כדור ענק לוחט של גז מימן והליום



רדיוס השמש = 100 רדיוסי ארץ

מסת השמש = 300,000 מסות ארץ

# הקווים הספקטרליים של אטום המימן

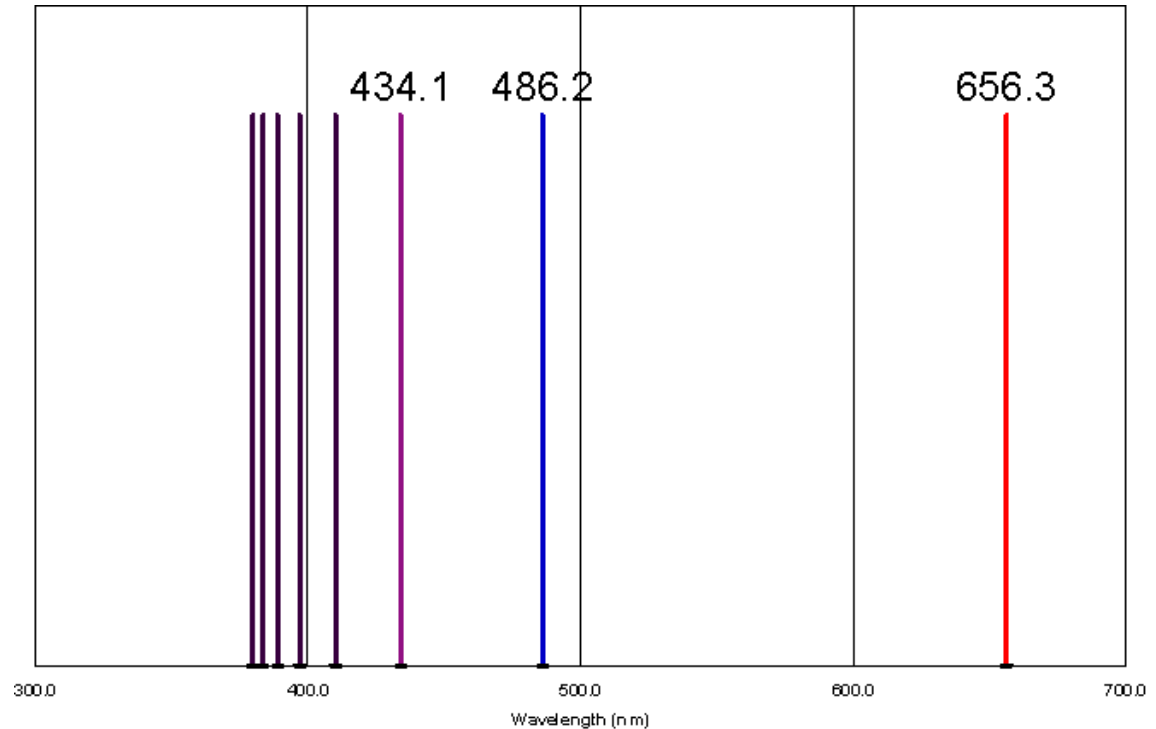
סדרת בלמר

$$H_{\alpha} = 6563 \text{ \AA}$$

$$H_{\beta} = 4861 \text{ \AA}$$

$$H_{\gamma} = 4340 \text{ \AA}$$

$$H_{\delta} = 4101 \text{ \AA}$$



$$H_{\infty} = 3646 \text{ \AA}$$

# הקווים הספקטראליים של אטום המימן

סדרת ליימן

$$L_{\alpha} = 1215 \text{ \AA}$$

$$L_{\beta} = 1025 \text{ \AA}$$

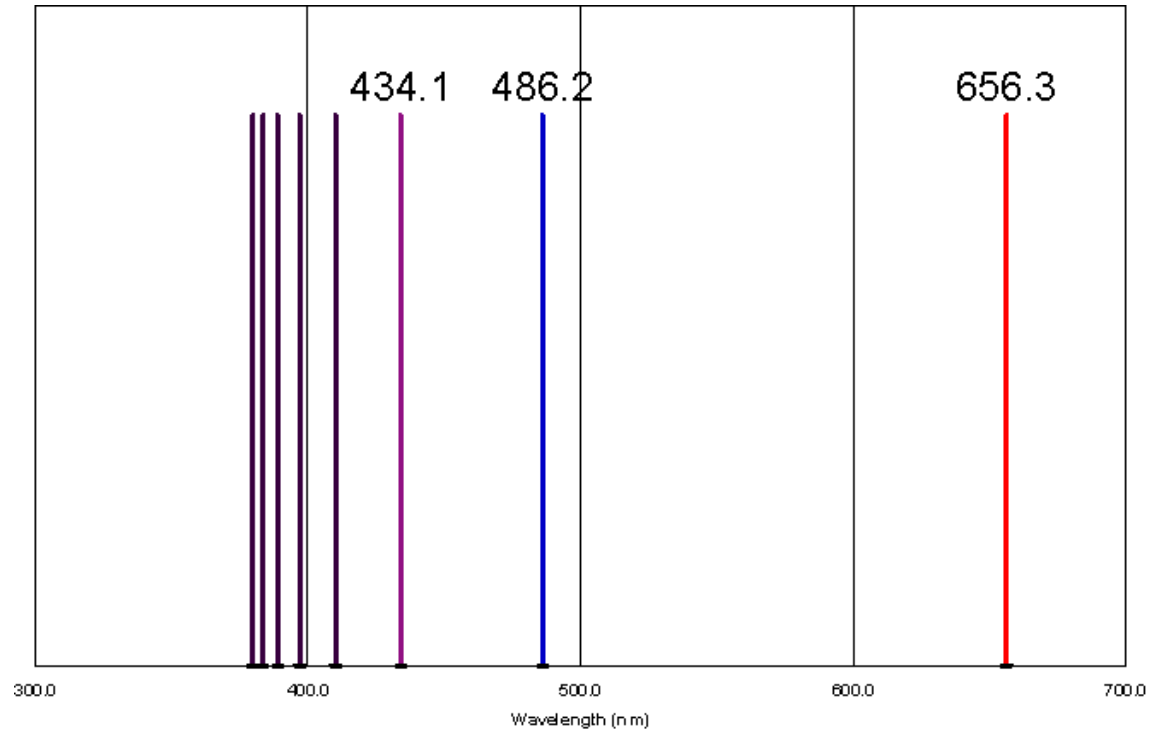
$$L_{\gamma} = 972 \text{ \AA}$$

$$L_{\infty} = 911.5 \text{ \AA}$$

# הקווים הספקטראליים של אטום המימן

סדרת בלמר

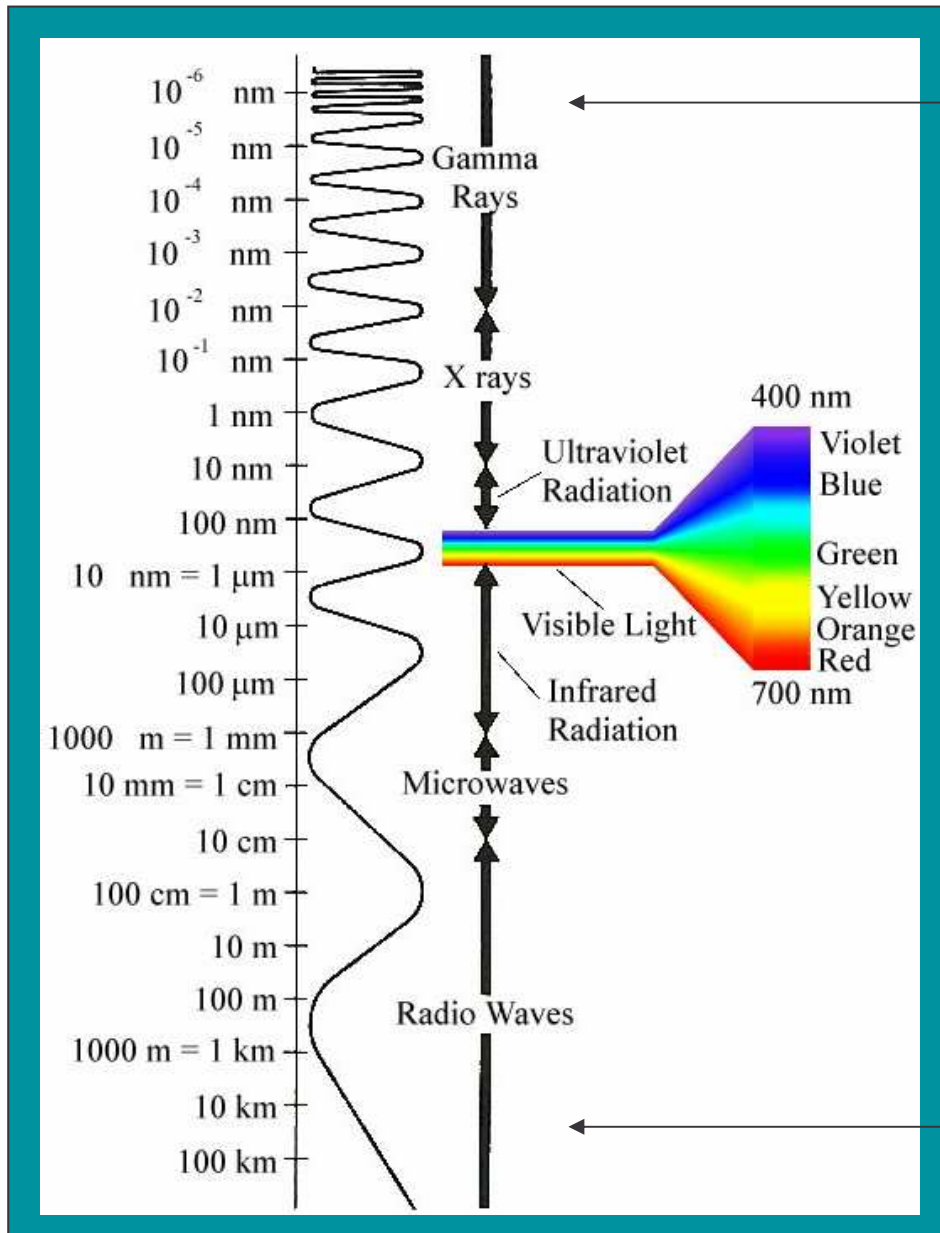
$$\frac{1}{\lambda_n} = A - \frac{B}{n^2}$$



$$\frac{1}{\lambda} = R \left( \frac{1}{2^2} - \frac{1}{n^2} \right)$$

$$R \approx 1.097 \times 10^5 \text{ cm}^{-1}$$

# The Electromagnetic Spectrum



Shortest wavelengths  
(Most energetic photons)

$$E = h\nu = hc/\lambda$$

$h = 6.6 \times 10^{-27}$  [erg\*sec]  
(Planck's constant)

Longest wavelengths  
(Least energetic photons)

# הקווים הספקטראליים של אטום המימן

$$\frac{1}{\lambda_n} = R \left( \frac{1}{1^2} - \frac{1}{n^2} \right) \text{Lyman}$$

$$\frac{1}{\lambda_n} = R \left( \frac{1}{2^2} - \frac{1}{n^2} \right) \text{Balmer}$$

$$\frac{1}{\lambda_n} = R \left( \frac{1}{3^2} - \frac{1}{n^2} \right) \text{Paschen}$$

$$\frac{1}{\lambda_n} = R \left( \frac{1}{4^2} - \frac{1}{n^2} \right) \text{Brackett}$$

# הקווים הספקטראליים של אטום המימן

$$\frac{1}{\lambda_{m,n}} = R \left( \frac{1}{m^2} - \frac{1}{n^2} \right)$$

$$E_{m,n} = \frac{hc}{\lambda_{m,n}} = hcR \left( \frac{1}{m^2} - \frac{1}{n^2} \right) = 13.6 \left( \frac{1}{m^2} - \frac{1}{n^2} \right) \text{ eV}$$

$$E_{m,n} = 13.6 \left( \frac{1}{m^2} - \frac{1}{n^2} \right) \text{ eV} \quad n > m$$



## רמות האנרגיה של אטום המימן

$$E_k = -13.6 \frac{1}{k^2} \text{ ev}$$

$$E_{m,n} = E_n - E_m$$

$$E_{m,n} = 13.6 \left( \frac{1}{m^2} - \frac{1}{n^2} \right) \text{ ev}$$

$$n > m$$

$$\Delta E_{\text{electron}} = \frac{hc}{\lambda} = hcR \left( \frac{1}{n^2} - \frac{1}{m^2} \right)$$
$$hcR = 13.6 \text{ eV}$$

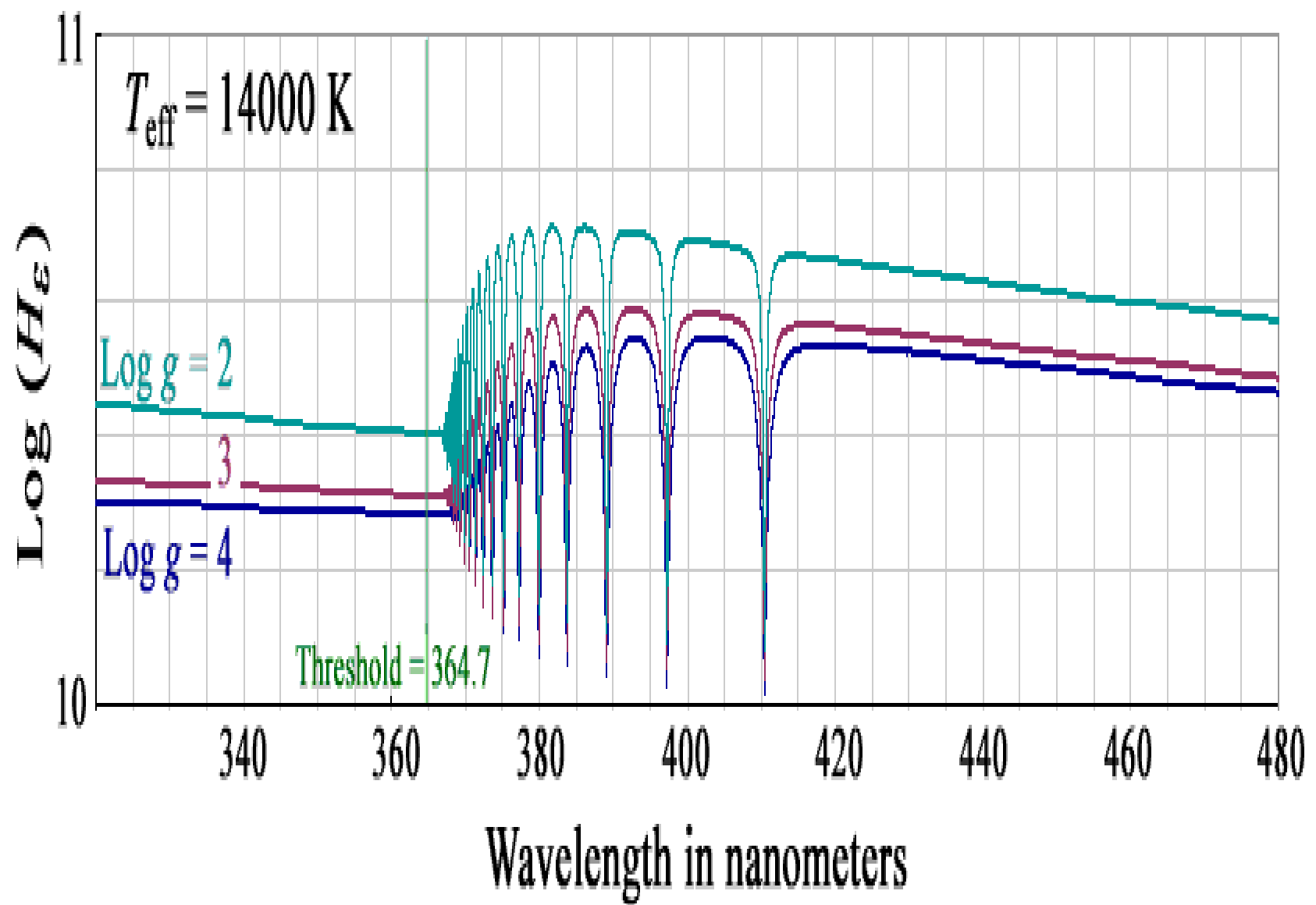
קוונטיזציה של רמות האנרגיה של אטום המימן

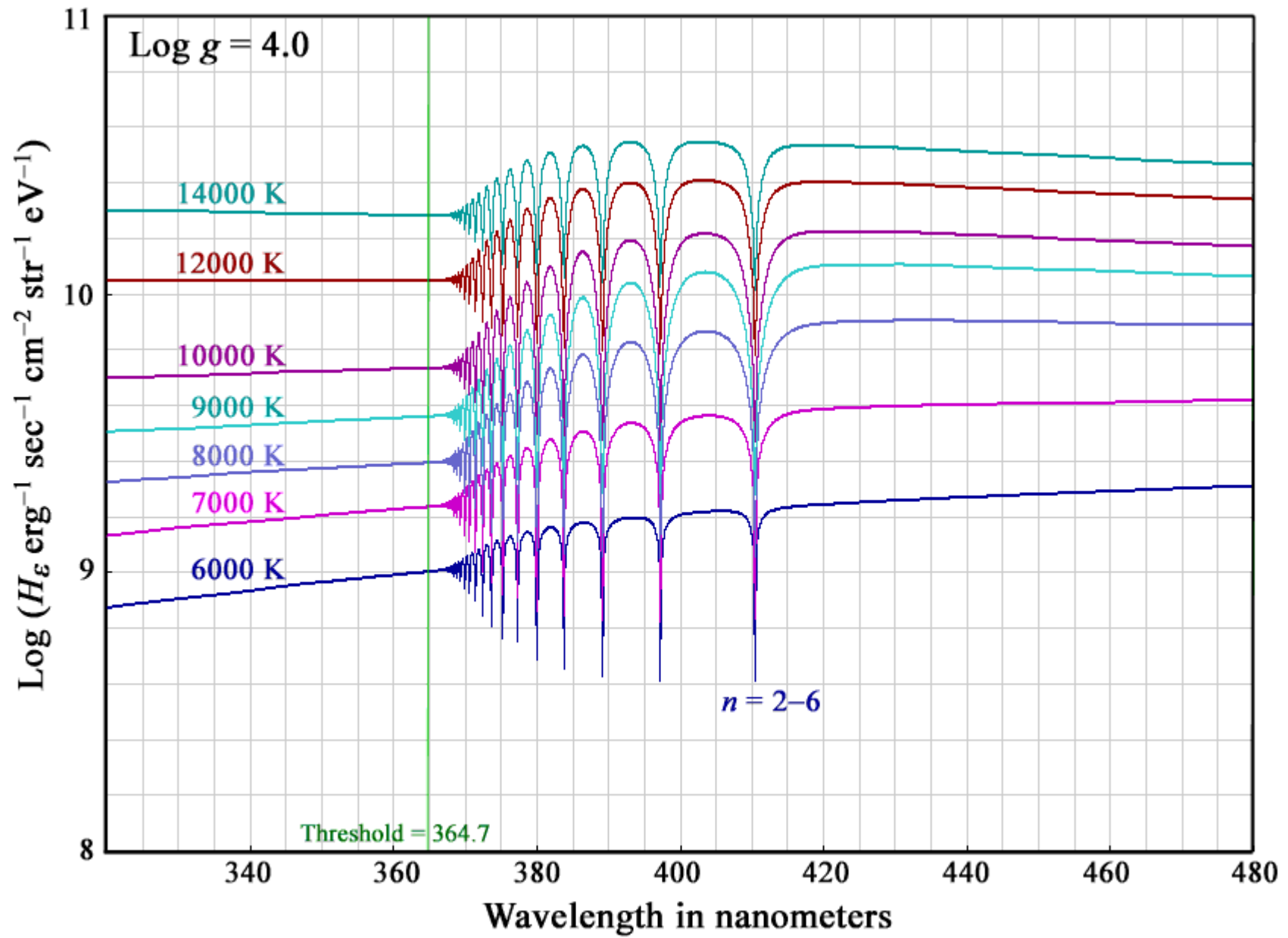
פליטה ספונטנית

בליעה פוטונים רק עם אנרגיה של מעבר בין רמות

# רמות האנרגיה של אטום המימן







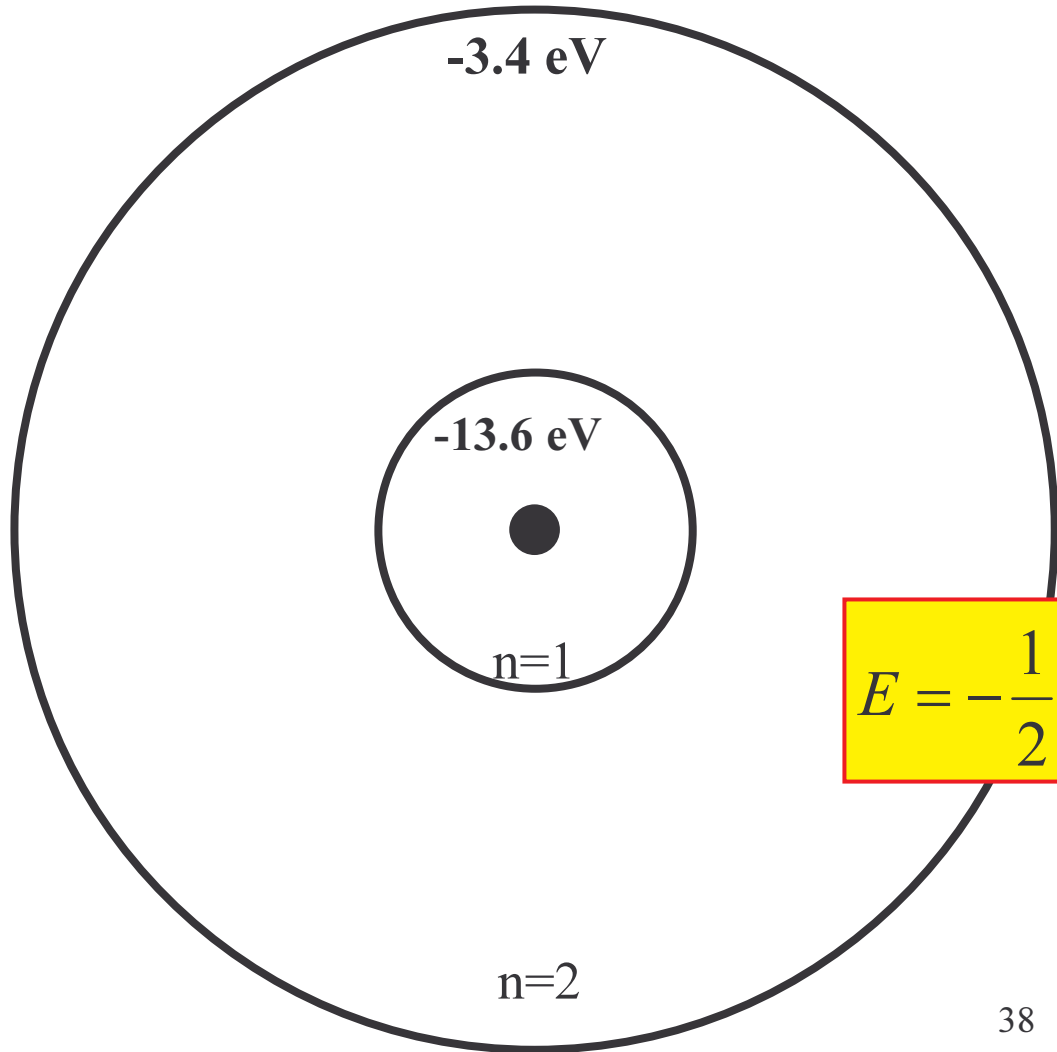
$$E_n = -13.6 \text{ eV} \frac{1}{n^2}$$

רמות האנרגיה של אטום המימן

$$m \frac{v^2}{r} = \frac{e^2}{r^2}$$

$$mv^2 = \frac{e^2}{r}$$

$$E_{kin} = \frac{1}{2} |U|$$

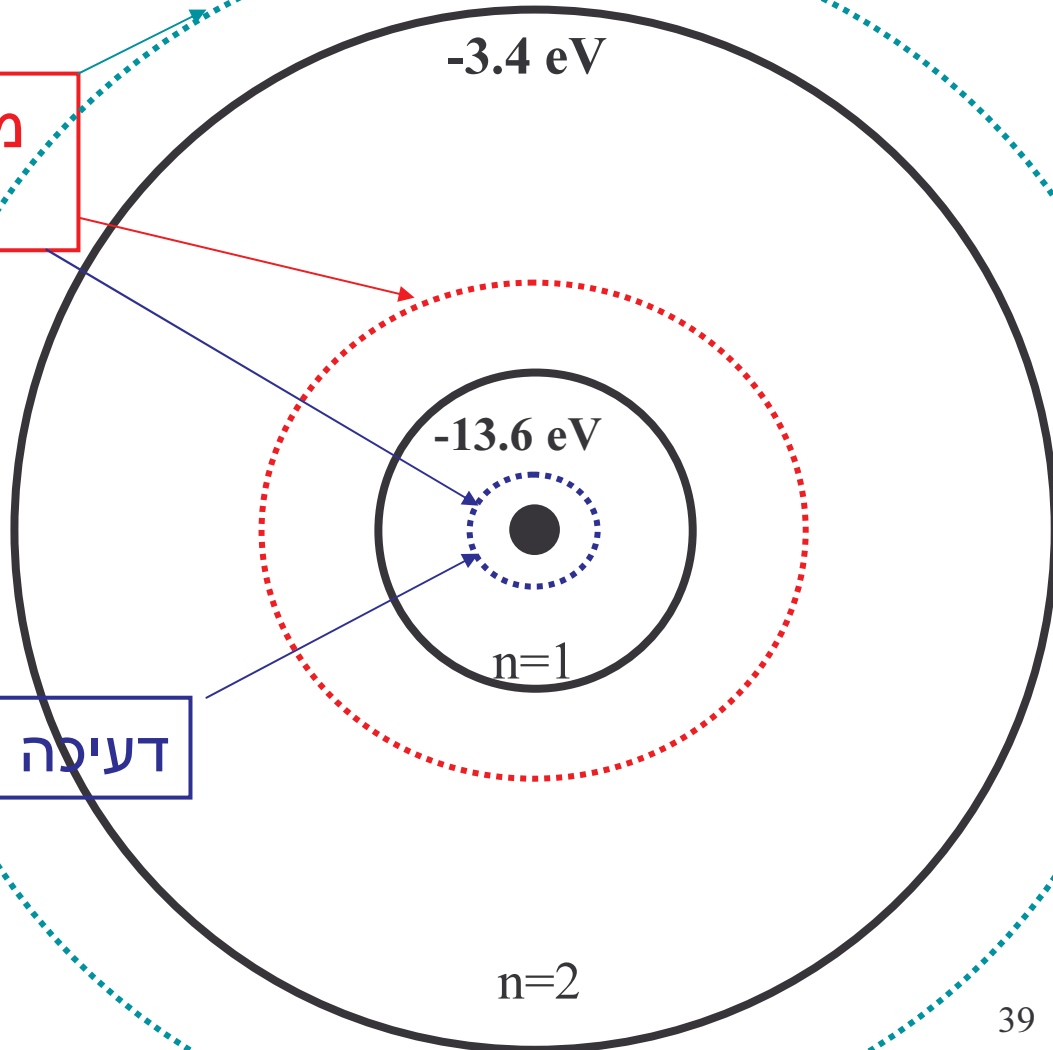


$$E = -\frac{1}{2} \frac{e^2}{r}$$

# המסלולים של אטום המימן

מדוע המסלולים האחרים  
אינם מותרים????

דעיכה של מטען שני במעגל



## Matter Waves ?

Maurice de Broglie: French experimental physicist, supports Compton's view of particle nature of radiation



Louis de Broglie: Maurice's brother, switches from graduate study in "History of physics", writes a revolutionary thesis



1924: He proposes that the wave-particle duality applies not only to radiation, but also to matter.

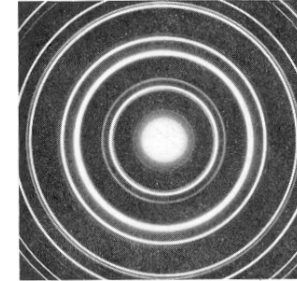
That is, just as a photon has light wave associated with it, an electron has an associated matter wave that governs its motion.

He further proposes that for matter and for radiation alike

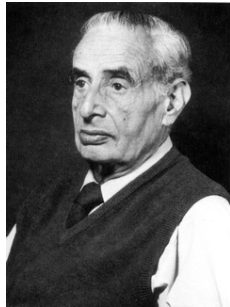
$$E = h\nu \quad \text{and} \quad p = h/\lambda \quad \left( \lambda = h/p \leftarrow \text{de Broglie relation} \right. \\ \left. \text{gives de Broglie wavelength} \right)$$



1921-1923: Davisson and Kunsman at Bell labs working on thermionic emission (actually studying electron emission from metal surfaces under electron bombardment) observe electron diffraction, but do not recognize it as such. (Davisson thinks pattern due to crystal structure of the metal).



(recall de Broglie paper wasn't until 1924)



1925: Walter Elsasser (who has read de Broglie's papers) explains Davisson-Kunsman results as electron diffraction – suggests it might be tested in same way as x-rays (scattering on a crystal)

1926: Davisson finally hears of de Broglie's ideas and devises (with Germer)



definitive experiment to test de Broglie hypothesis:

more on Elsasser:

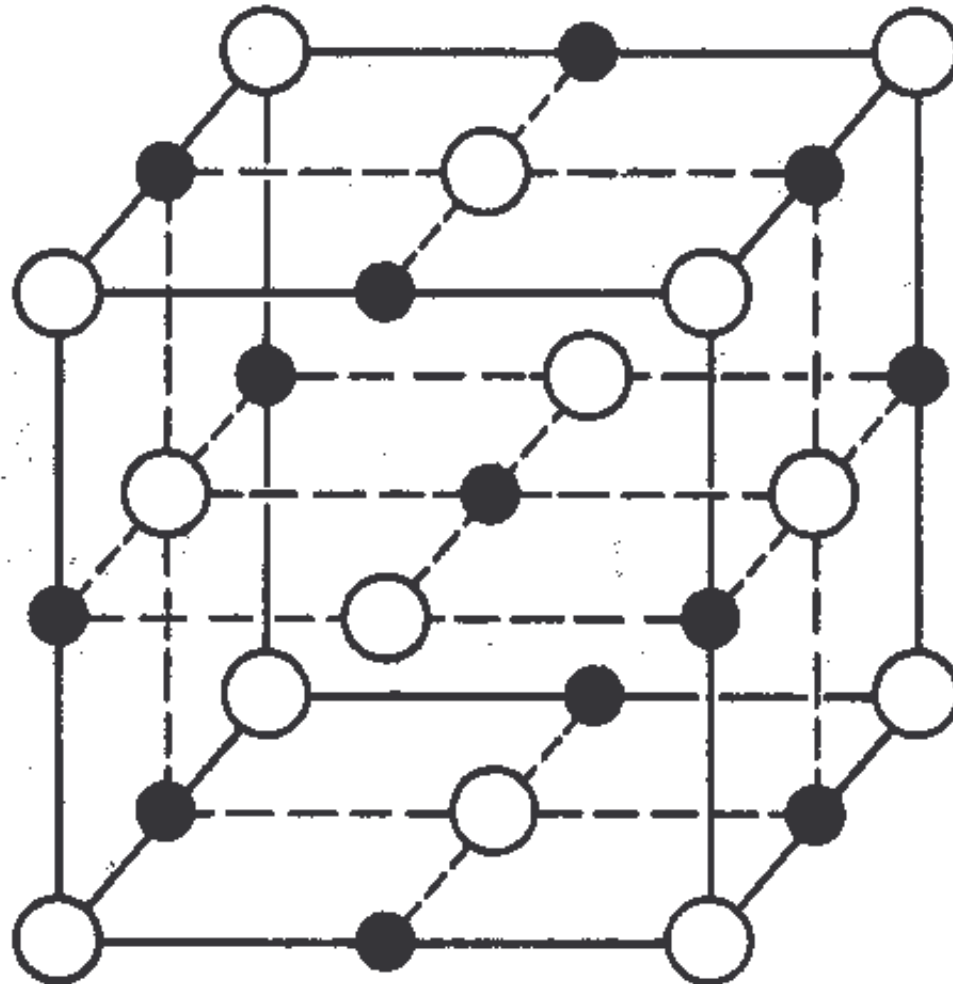
First to recognize the shell structure of atomic nuclei.

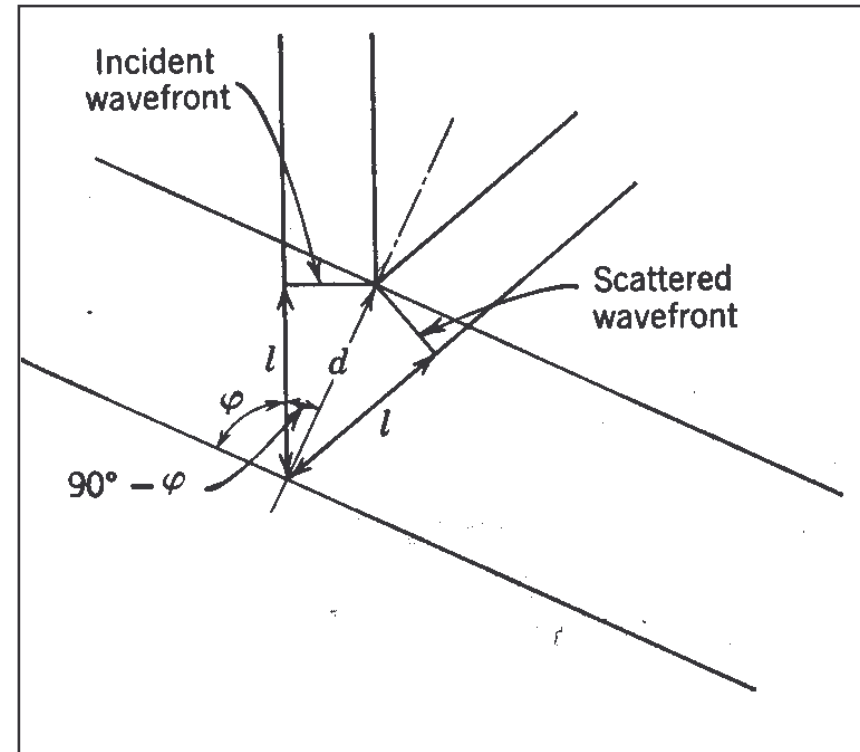
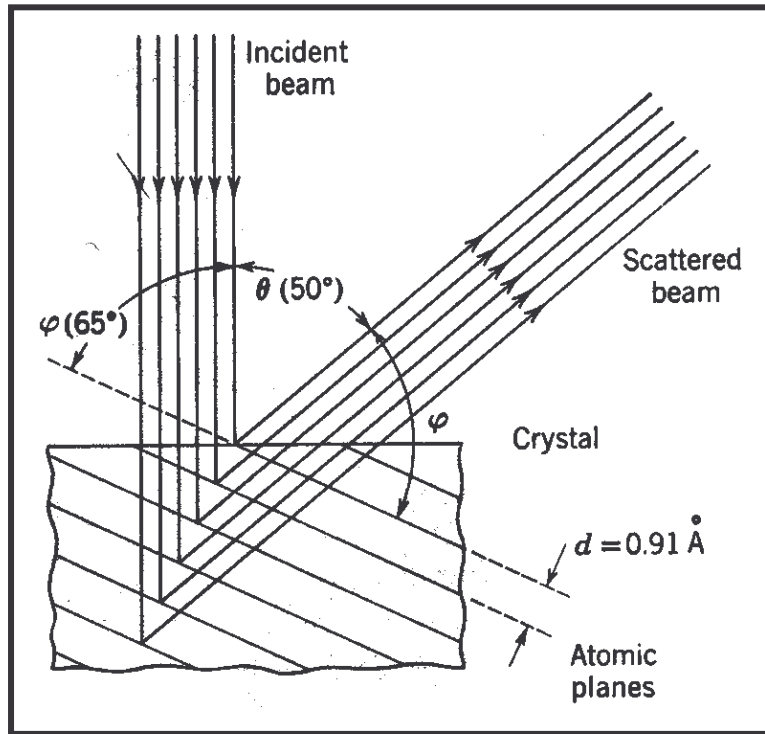
geophysics: provided important insights about the radiative transfer of heat in the atmosphere; fathered the generally accepted dynamo theory of the earth's magnetism.

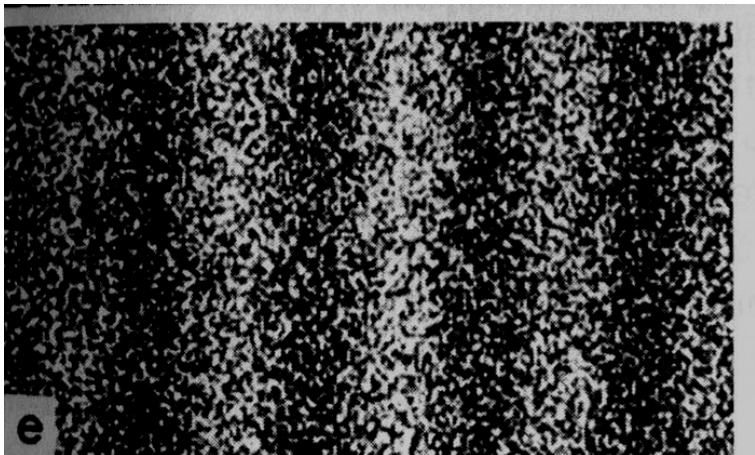
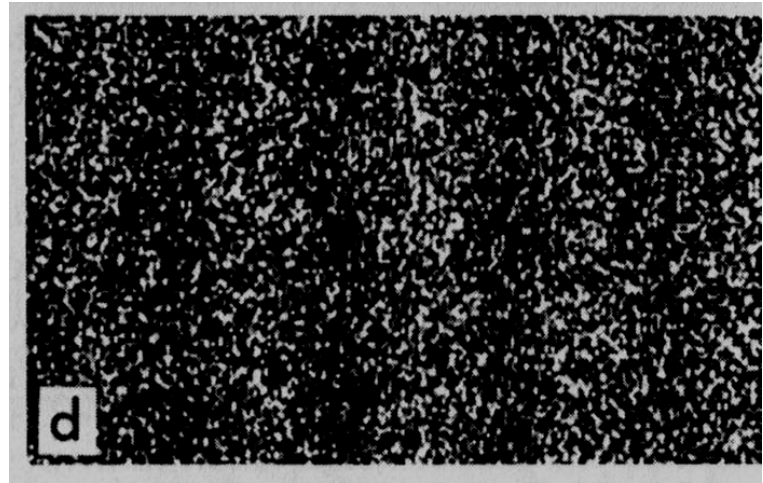
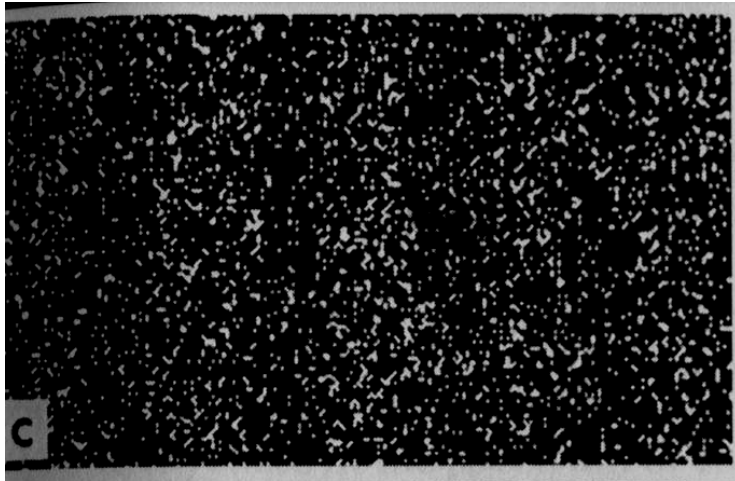
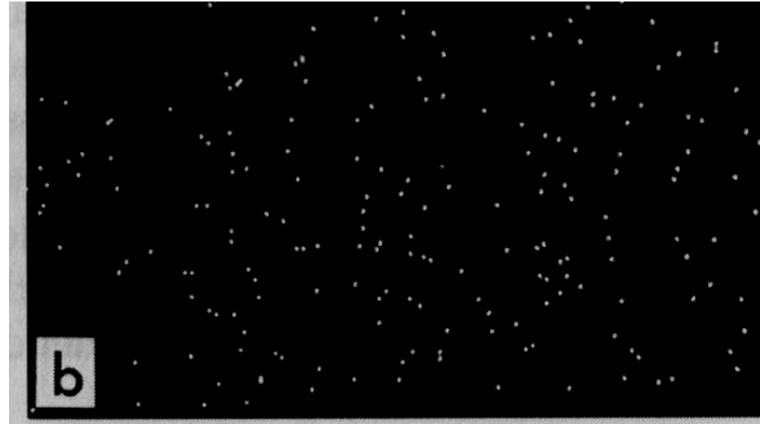
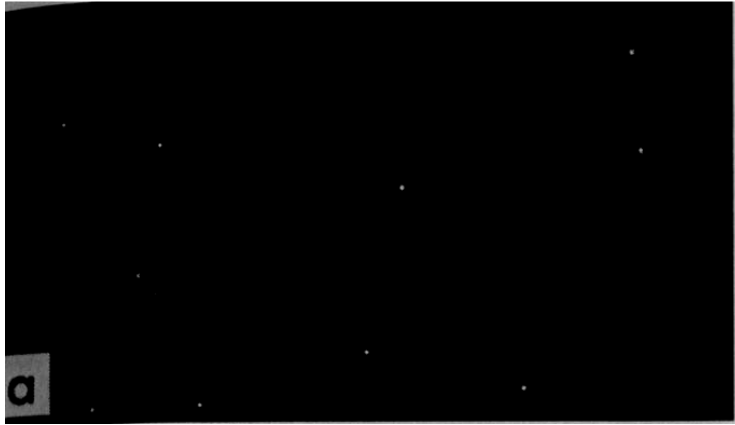
biophysics: last fifty years of his life spent developing a theory of organisms

Davisson-Germer apparatus:

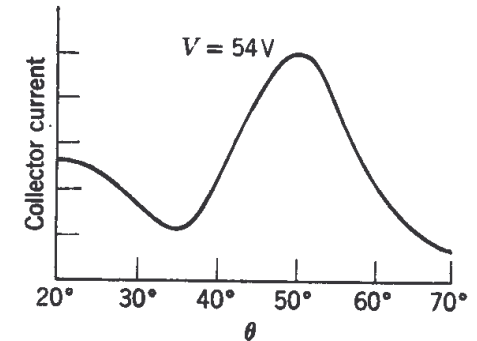
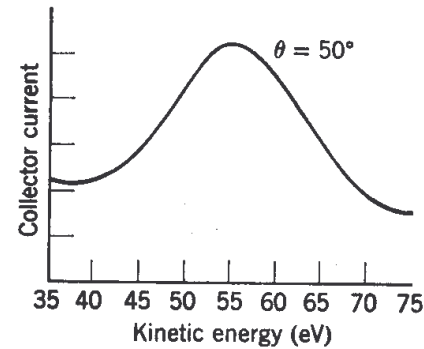
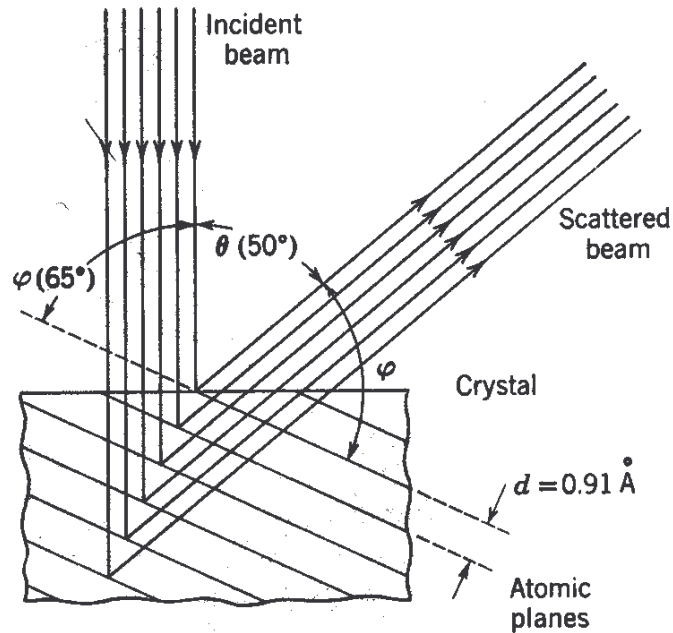








This figure is for 54 eV electrons at 50° for which we have a peak:



What wavelength do we measure ?  
 (recall  $n\lambda = 2d\sin\phi$ )  $\lambda = 1.65 \text{ \AA}$

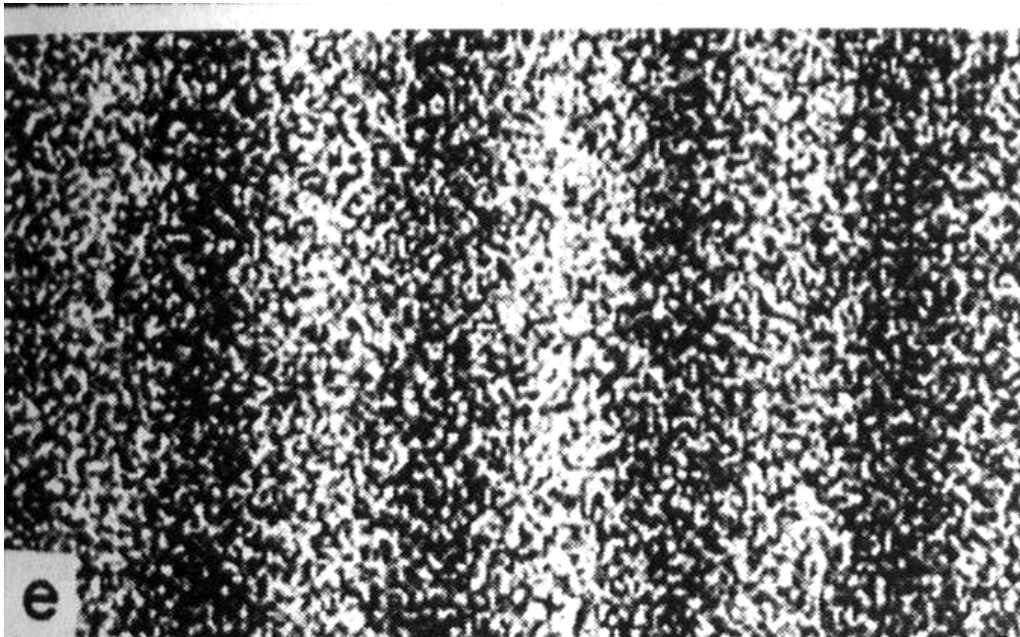
What de Broglie wavelength do we expect for 54 eV electrons ?

$\lambda = h/p = 1.65 \text{ \AA}$  (using  $p$  value from Example 3-1)

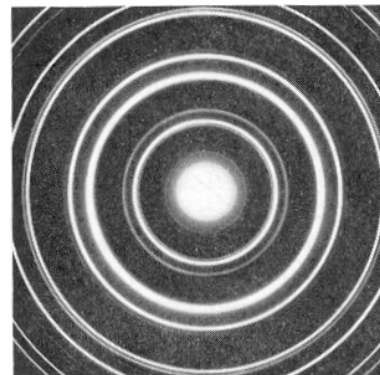
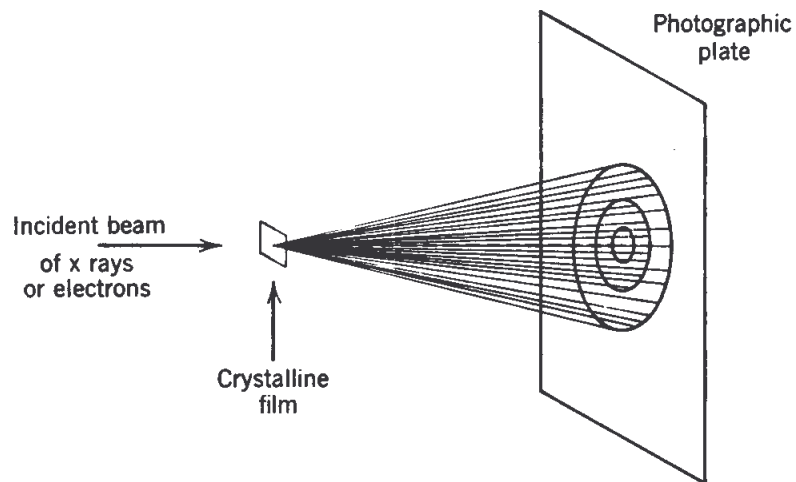
Let's look more closely at the dependence on electron energy ( $eV_0$ ).

Start by asking: What is the relationship between  $\lambda$  and  $V_0$  ?





1927: G. P. Thompson (son of JJ) sends a beam of higher energy electrons through a thin crystalline film



Where are all these lines coming from ?

And why is there such a variation in intensity ?

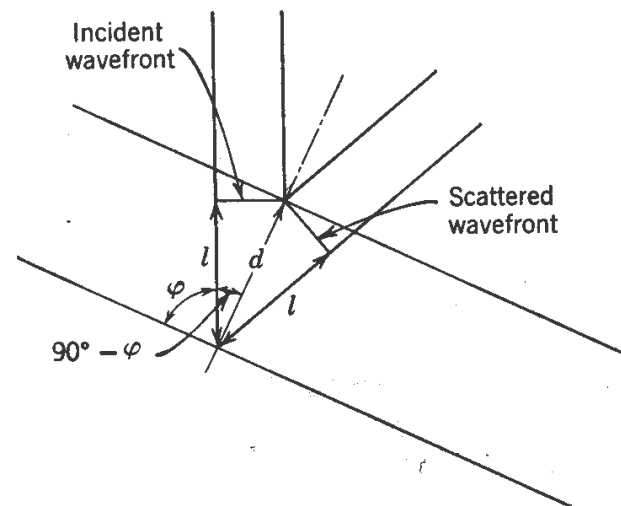
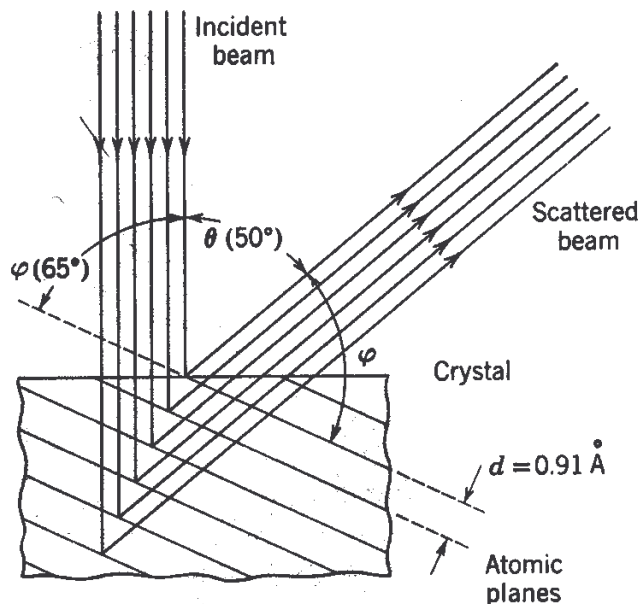


**Key point: Interference is NOT between waves associated with one electron and waves associated with another. Interference is between different parts of the wave associated with a SINGLE electron that have been scattered from various regions of the crystal.**

How would you verify this experimentally ?

Reduce electron beam intensity until it is so low that electrons are going through one at a time.

process is analogous to Bragg scattering of x-rays

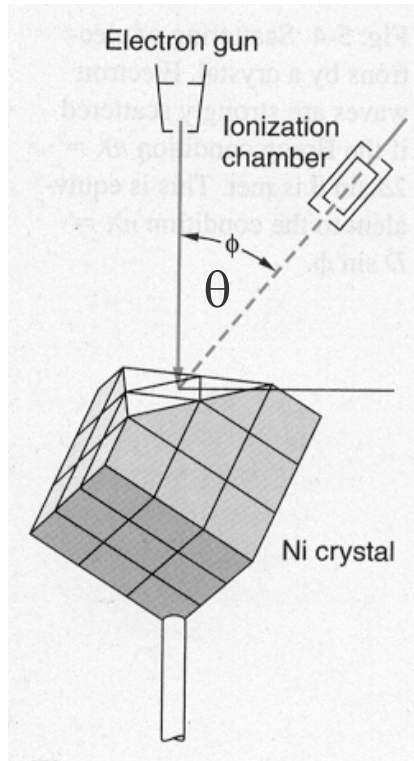


constructive interference if

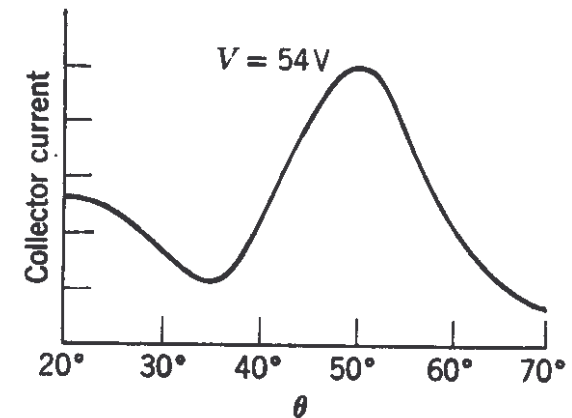
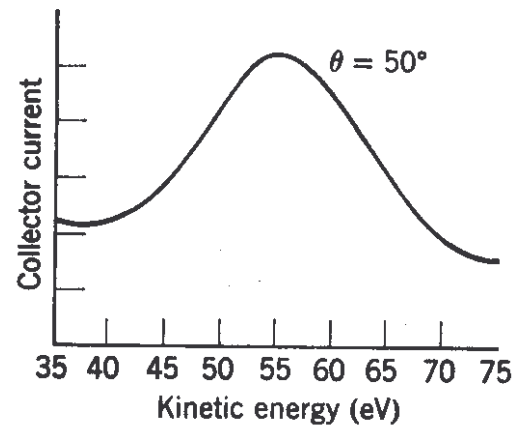
$$n\lambda = 2d\sin\phi$$

(Bragg relation)

## Davisson-Germer apparatus:



Look at intensity as a function of electron kinetic energy (eV) and  $\theta$



How do we explain these peaks ?

Constructive interference of waves scattered by the periodic arrangement of atoms into planes of the crystal

Demonstration of de Broglie's hypothesis!

Send matter particles through small aperture and look for diffraction ( $\lambda/a \sim 1$ )

What kind of matter particles would be best ?

A particle with the largest possible wavelength, baseball is no good, maybe electron

Clearly, the smaller the aperture, the better – back in the 1920s, the smallest aperture available to experimenters was  $\sim 1 \text{ \AA}$ .

Any idea how they achieve such small apertures ?

Spacing between adjacent planes of atoms in a solid

